

# NEUTRINOLESS DOUBLE BETA DECAY WITH SNO+

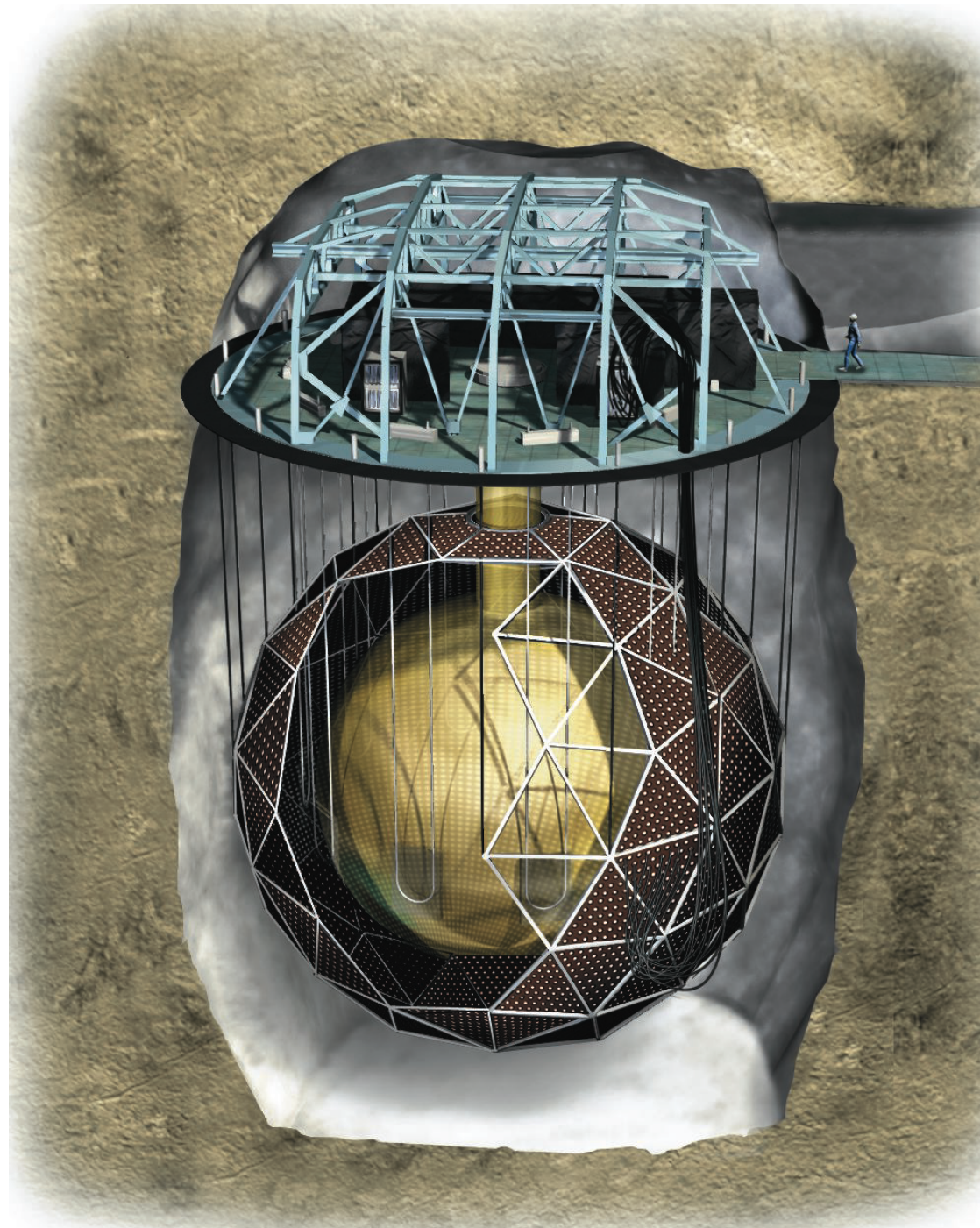
CAP Congress 2013, Montreal  
May 30<sup>th</sup>, 2013

Alex Wright  
IPP/Queen's University  
For the SNO+ Collaboration





- SNO heavy water replaced by 780 tonnes of liquid scintillator
- ~9500 PMTs
- 1500 + 5300 tons ultra-pure water shielding
- New rope net to hold down the 6m radius acrylic vessel
- 6800' underground in SNOLAB



# The SNO+ Collaboration

ARMSTRONG ATLANTIC STATE UNIVERSITY 

BLACK HILLS STATE UNIVERSITY 

BROOKHAVEN NATIONAL LABORATORY 

LAURENTIAN UNIVERSITY 

LIP COIMBRA 

LIP LISBOA 

OXFORD UNIVERSITY 

QUEEN MARY, UNIVERSITY OF LONDON 

QUEEN'S UNIVERSITY 

SNOLAB 

TECHNICAL UNIVERSITY OF DRESDEN 

TRIUMF 

UNIVERSITY OF ALBERTA 

UNIVERSITY OF CALIFORNIA – BERKELEY   
& LAWRENCE BERKELEY NATIONAL LABORATORY 

UNIVERSITY OF CHICAGO 

UNIVERSITY OF LIVERPOOL 

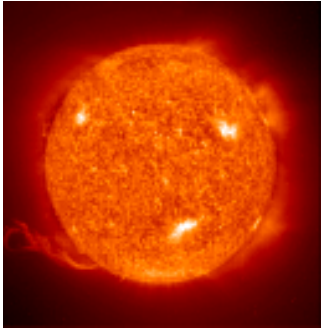
UNIVERSITY OF NORTH CAROLINA AT CHAPEL HILL 

UNIVERSITY OF PENNSYLVANIA 

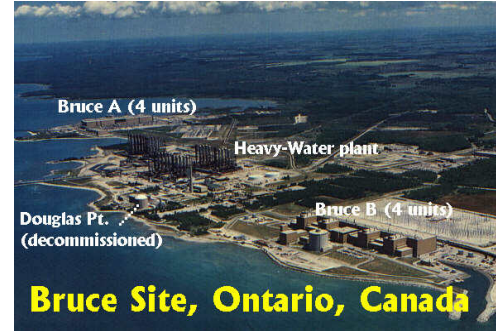
UNIVERSITY OF SUSSEX 

UNIVERSITY OF WASHINGTON 

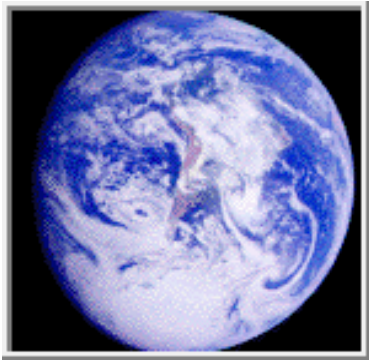
# SNO+ Physics



**Low Energy Solar Neutrinos**



**Reactor Antineutrinos**

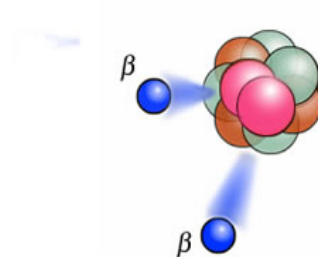
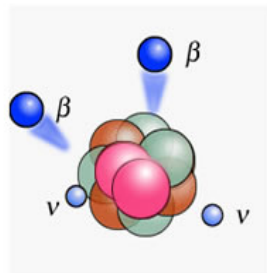


**Geo-Neutrinos**

**Supernova Neutrinos**



**Neutrinoless Double Beta Decay**

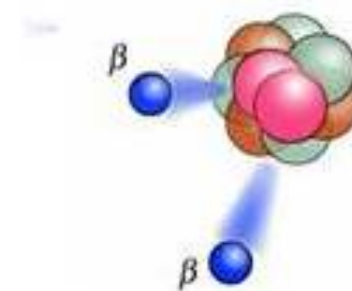
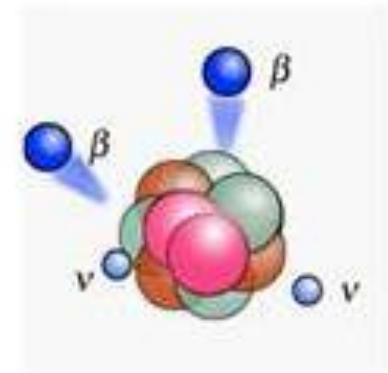




# Neutrinoless Double Beta Decay

- Are neutrinos Majorana or Dirac particles?
    - Are they their own anti-particles?
  - In double beta decay, a nucleus releases two electrons and two antineutrinos:
- $(A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\bar{\nu}_e$
- If neutrinos are Majorana, sometimes neutrinoless double beta decay occurs:

$$(A, Z) \rightarrow (A, Z + 2) + 2e^-$$



**Detection of neutrinoless double beta decay proves that neutrinos are Majorana and provides information about the neutrino mass.**

Searching for neutrinoless double beta decay involves looking for a tiny monoenergetic peak at the end of a large double beta decay continuum.

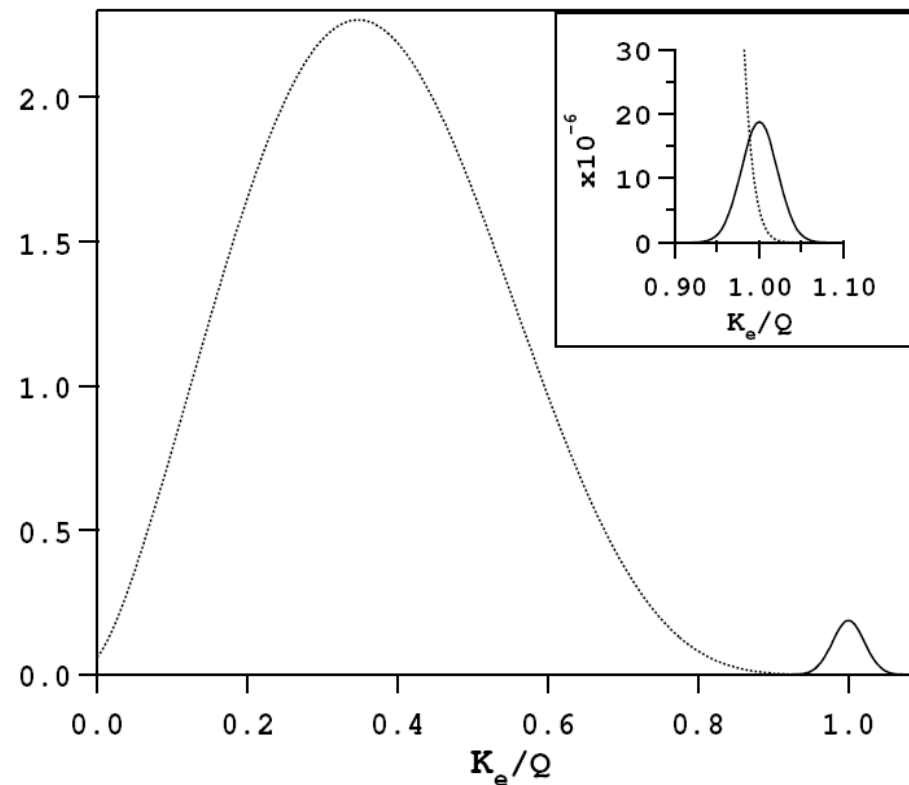


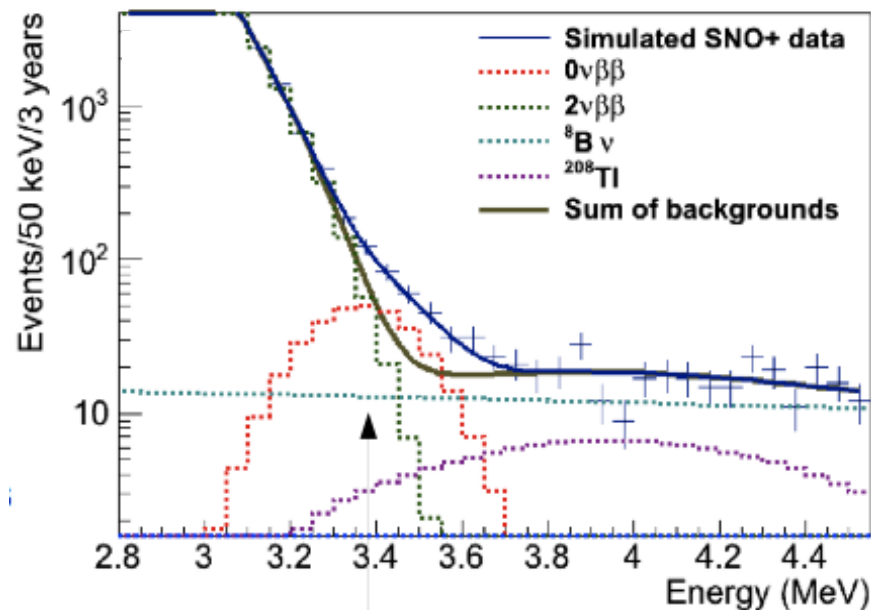
Image from Elliott and Vogel, hep-ph/0202254

***D.B.D. experiments need good energy resolution, low backgrounds, and large amounts of isotope.***

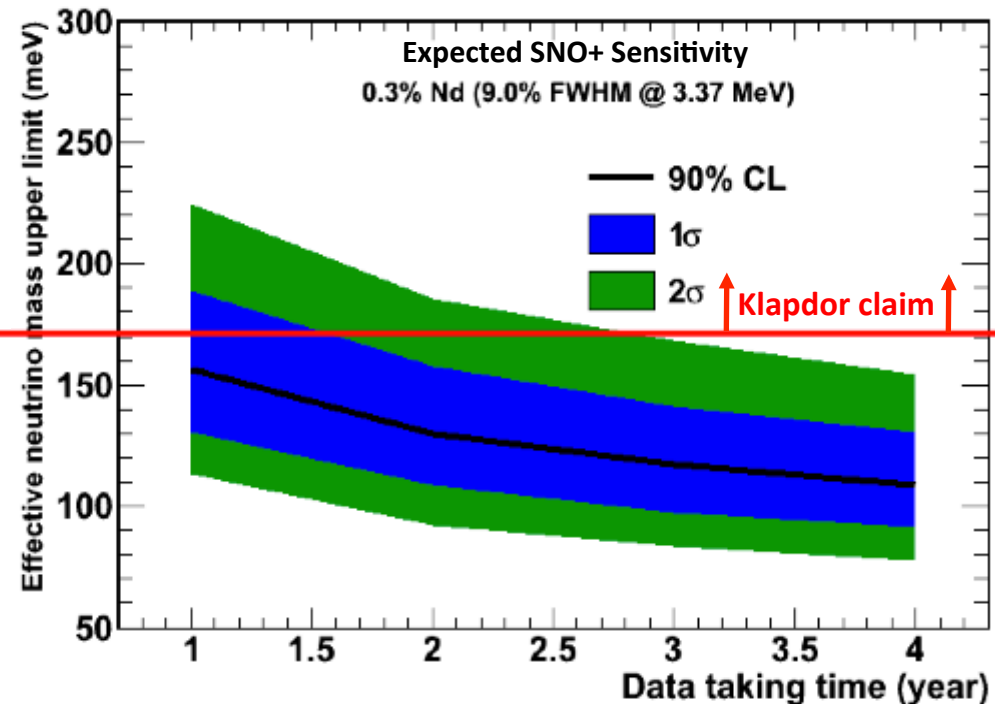


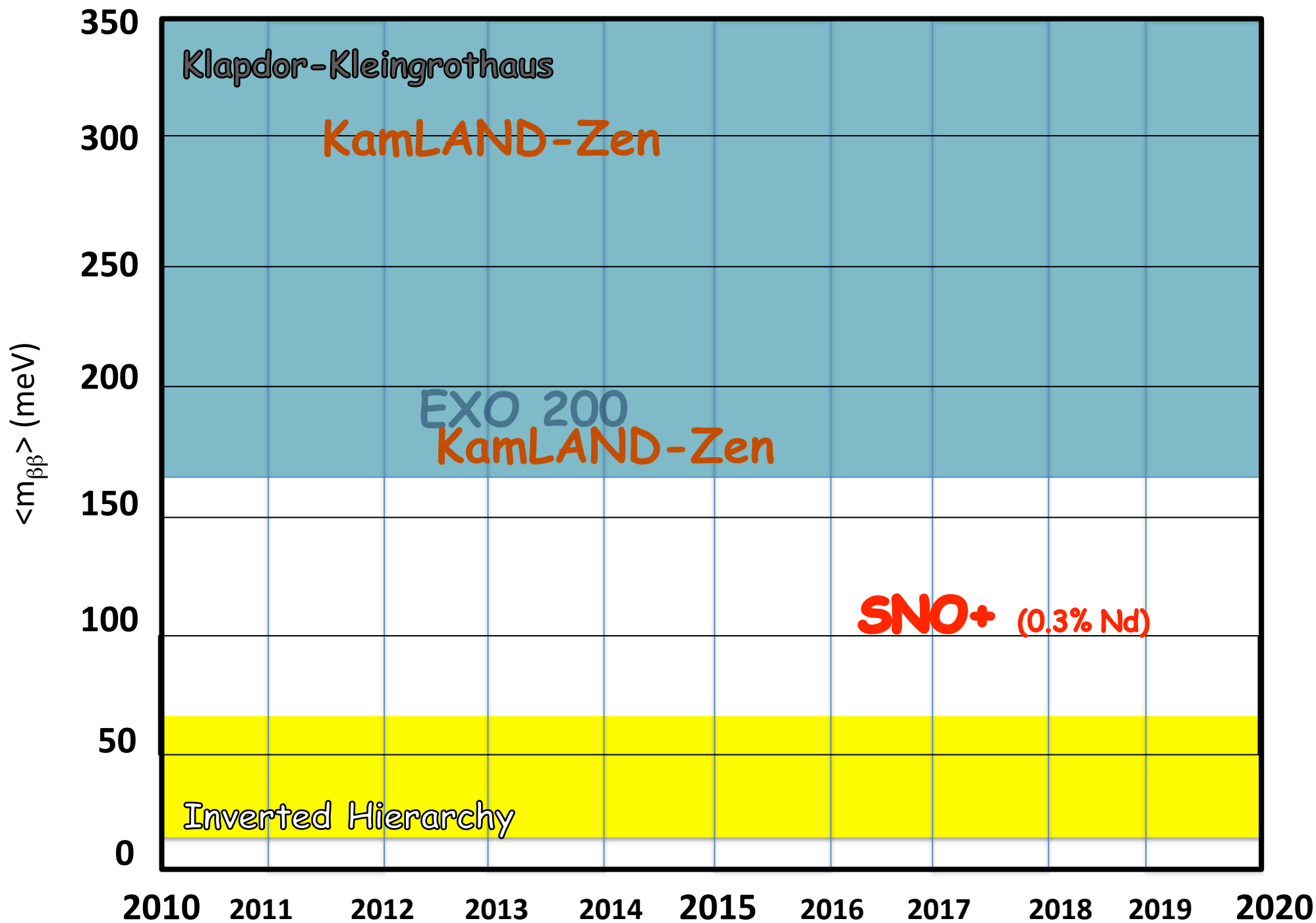
# Neutrinoless Double Beta Decay in SNO+

Loading neodymium into the SNO+ scintillator gives 140kg  $^{150}\text{Nd}$  at 0.3% loading of natural neodymium (limited by optics).



- Simulated SNO+ data
- Signal at Klapdor level
  - 2.4 live-years of data
  - 50% fiducial volume
  - Borexino-level backgrounds
    - $^{214}\text{Bi}$  99.9% rejection
    - $^{208}\text{Tl}$  90% rejection





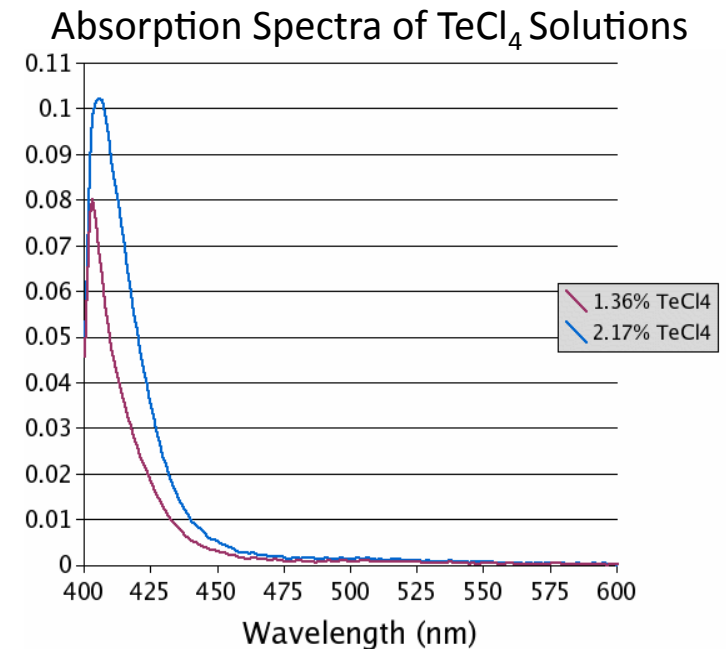
*To improve the sensitivity of the SNO+ measurement, need more isotope.*



# $\beta\beta$ Isotopes with Q-values >2 MeV

isotope	Q-value [MeV]	natural abundance
$^{48}\text{Ca}$	4.27	0.19%
$^{150}\text{Nd}$	3.37	5.6%
$^{96}\text{Zr}$	3.35	2.8%
$^{100}\text{Mo}$	3.03	9.6%
$^{82}\text{Se}$	3.00	9.2%
$^{116}\text{Cd}$	2.80	7.5%
$^{130}\text{Te}$	2.53	34.1%
$^{136}\text{Xe}$	2.48	8.9%
$^{124}\text{Sn}$	2.29	5.6%
$^{76}\text{Ge}$	2.04	7.8%
$^{110}\text{Pd}$	2.01	11.8%

$^{130}\text{Te}$  has highest natural abundance: *0.3% loading in SNO+ gives ~800kg of  $^{130}\text{Te}$  isotope*

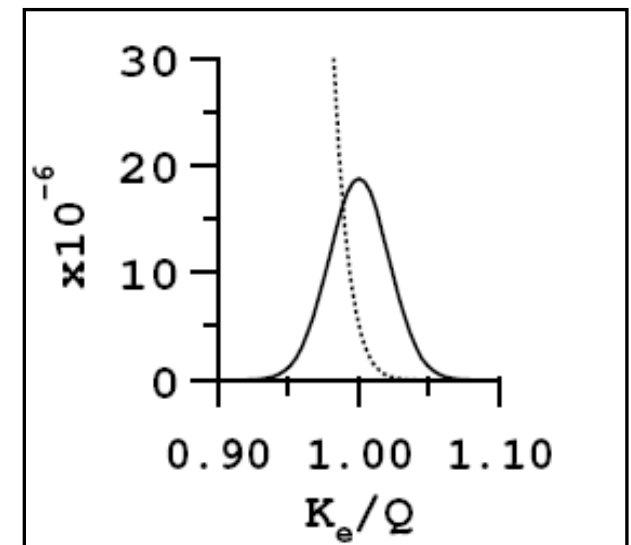


*Te has no absorption lines, so no a priori limit on loading concentration...*

# $\beta\beta$ Isotopes with Q-values $>2$ MeV

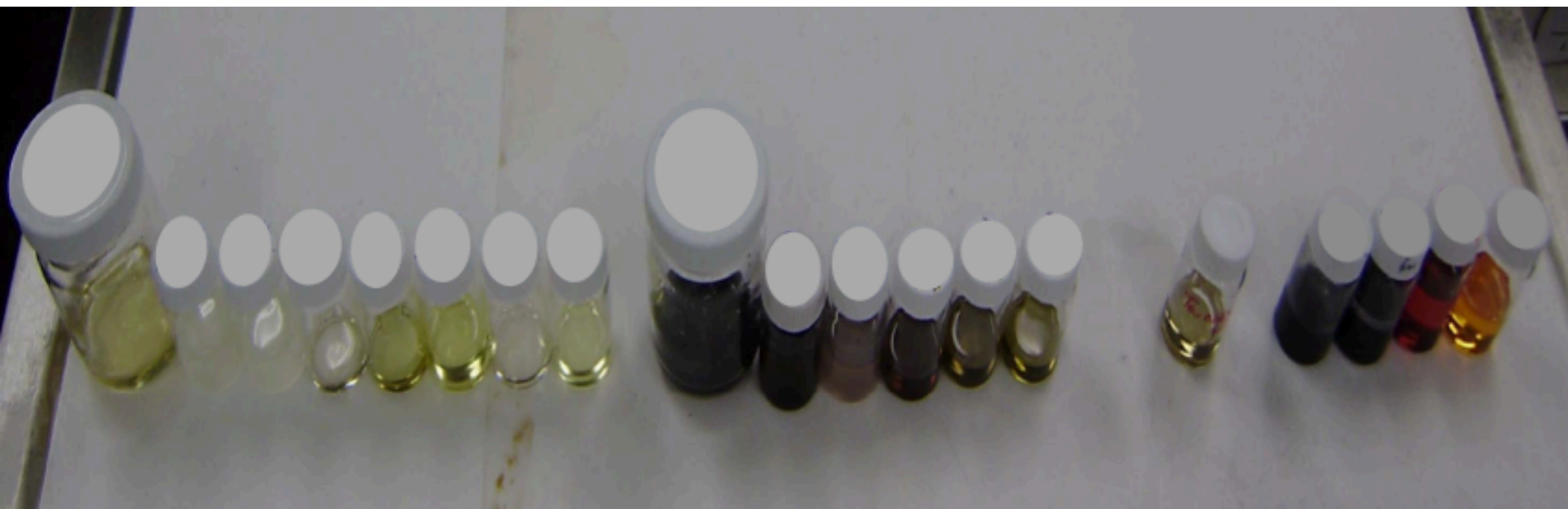
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In addition, the  $0\nu\beta\beta/2\nu\beta\beta$  ratio is predicted to be about 100x higher in Te than Nd – this is an advantage given the limited energy resolution in scintillator.

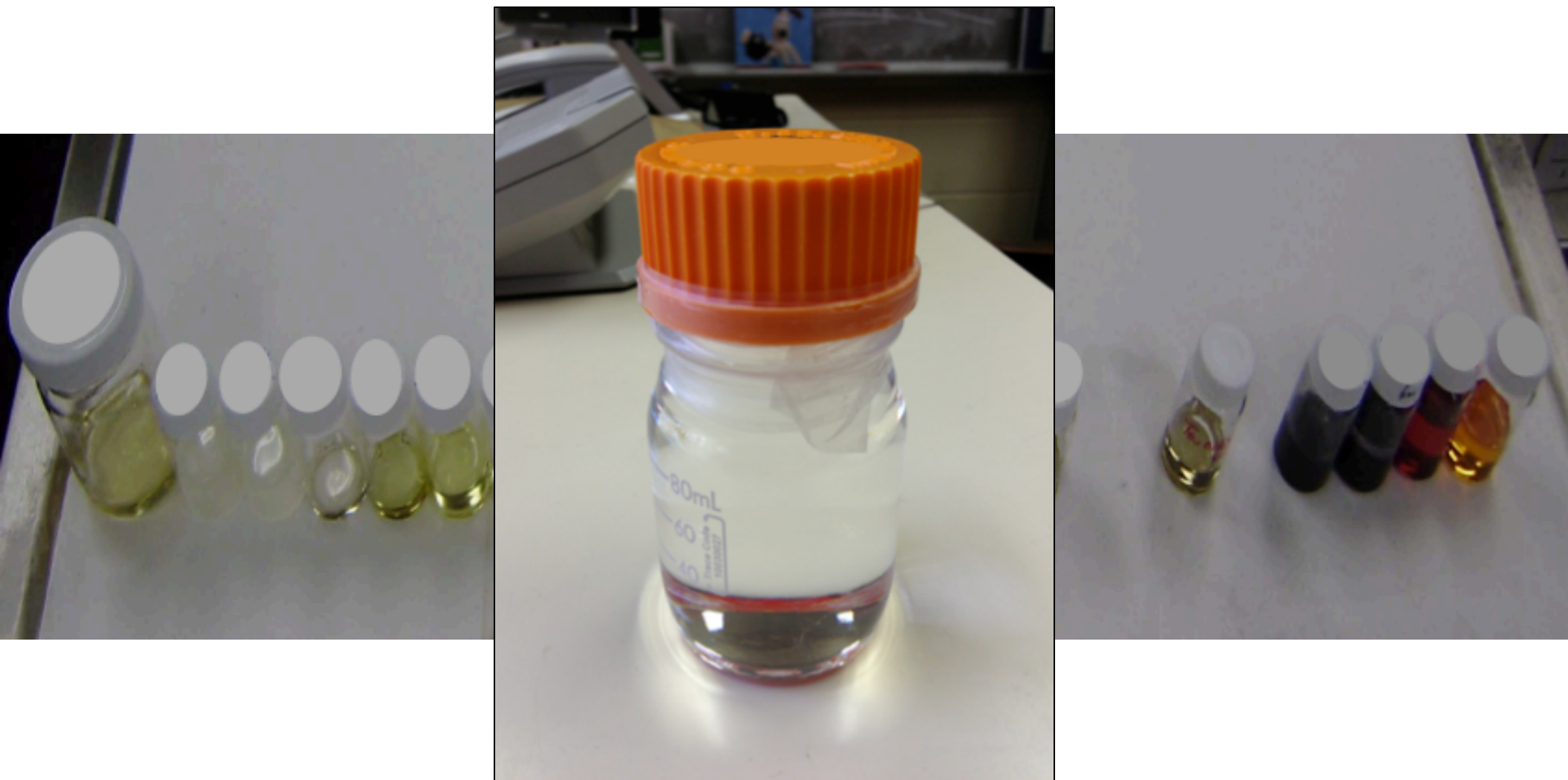




# Te-Loaded “Scintillators”



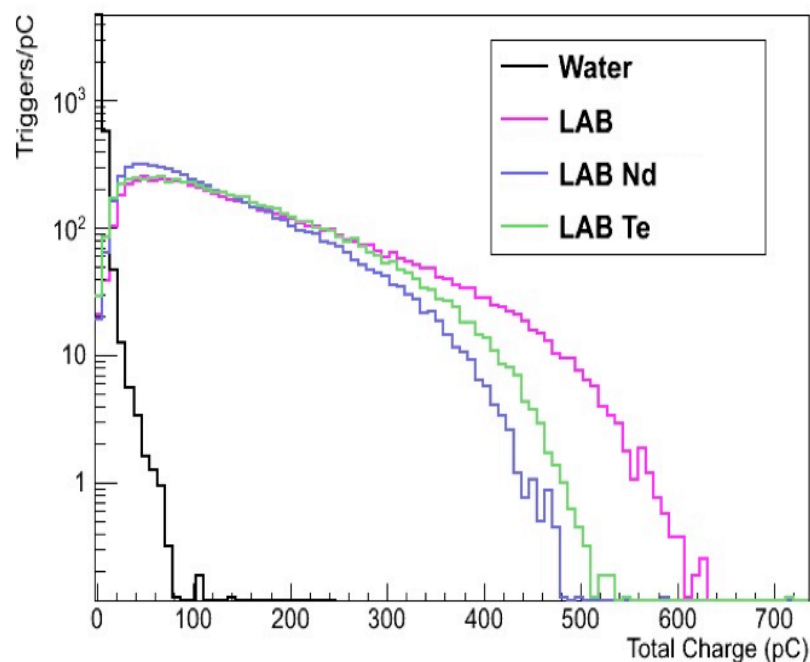
# Te-Loaded Scintillator



*SNO+ has developed stable Te-loaded LAB with good optical properties at percent-level loading.*

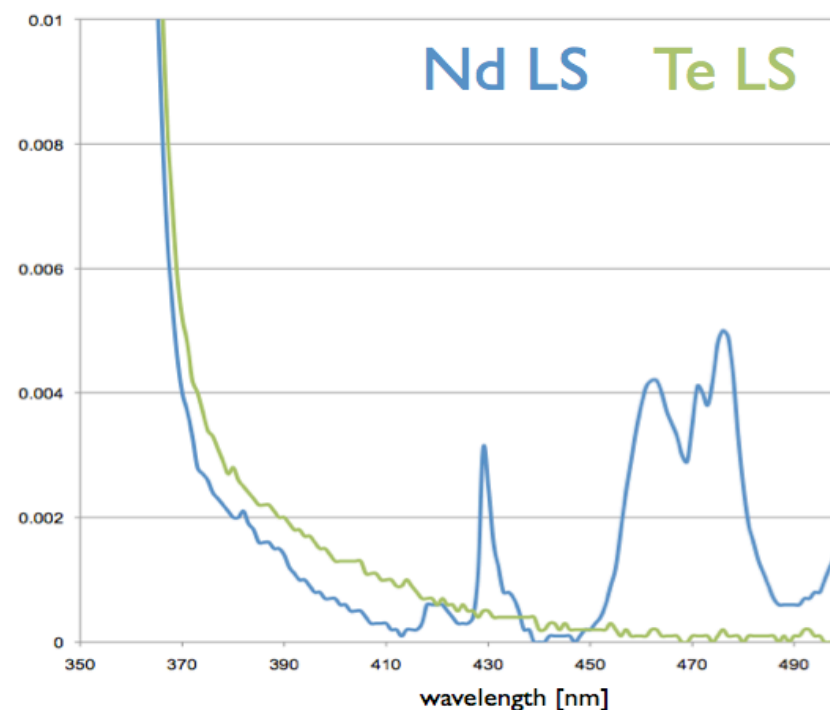


# Te-Loaded Scintillator



← Intrinsic light yield of Te-LS is good (better than Nd-LS)

- No absorption lines means that the light can be shifted out to 450-500nm
- Absorption may be further reduced by improvements in loading technique

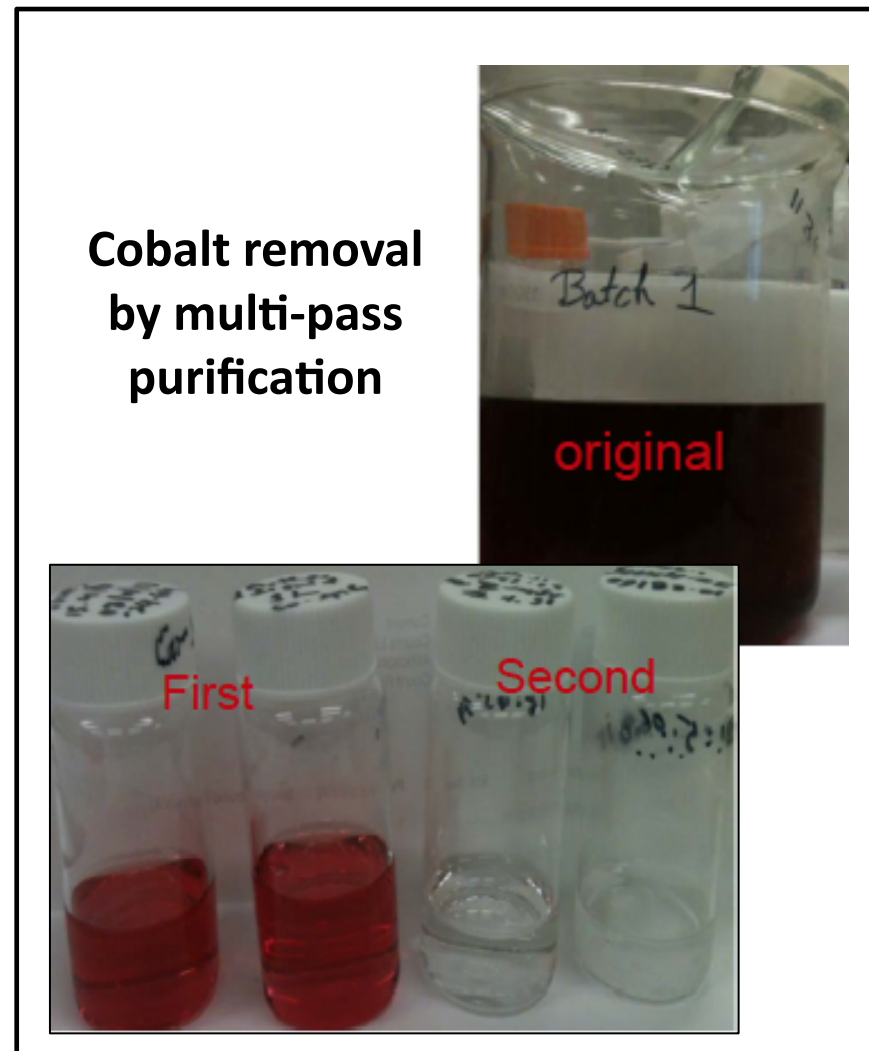


# Backgrounds

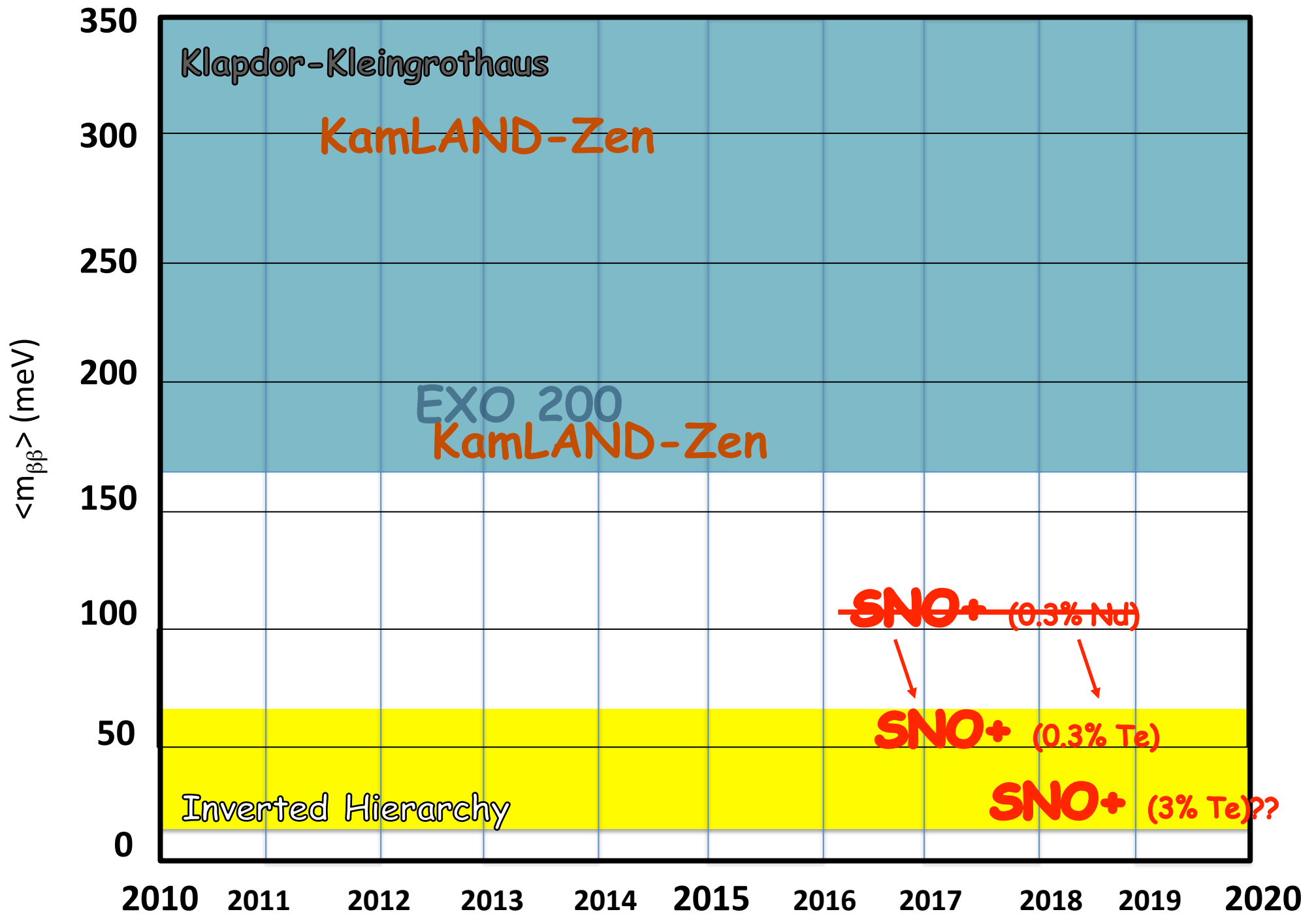
- Lower energy (2.53 MeV) endpoint makes Te susceptible to more radiogenic backgrounds than Nd
- 2.6 MeV gamma from external  $^{208}\text{Tl}$  suppressed by fiducialization
  - Internal  $^{208}\text{Tl}$  at higher energy
- Te endpoint overlaps with  $^{214}\text{Bi}$  spectrum ( $^{238}\text{U}$  chain)
  - U-chain backgrounds in liquid scintillator can be extremely low (<2 decays/day/100T)
  - $^{214}\text{Bi}$  can be suppressed by more than a factor of 1000 using the  $164\mu\text{s}$   $^{214}\text{Bi} - ^{214}\text{Po}$  delayed coincidence
  - SNO+ has developed techniques to purify Te to acceptable U/Th levels
  - Working to develop purification of other required chemicals

# Backgrounds

- A number of isotopes that are cosmogenically produced from Te have longish half-lives and have decays that overlap the  $0\nu\beta\beta$  energy region ( $^{214}\text{Sb}$ ,  $^{126}\text{Sn}$ ,  $^{88}\text{Y}$ ,  $^{110}\text{Ag}$ , etc...)
  - Effectively removed by our purification technique
  - Regrowth of cosmogenics requires at least a “polishing” purification underground





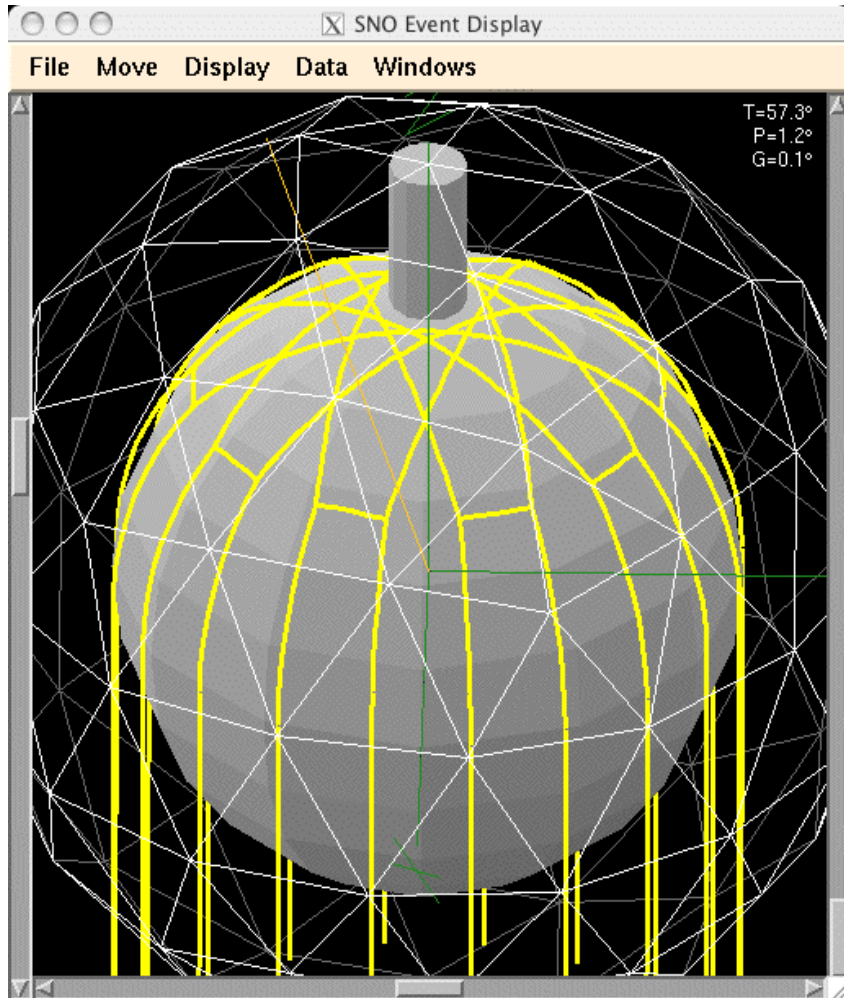


*SNO+ will pursue Te deployment for double beta decay, starting with 0.3% target loading.*

# SNO+ Status and Schedule

- Current: Construction phase
  - Install AV hold-down net
  - Upgrade electronics/DAQ
  - Install scintillator purification system
  - Upgrade calibration/covergas system
  - Clean Acrylic Vessel
- Summer 2013: Begin water fill
  - Buoyant test of hold-down net
  - Study backgrounds and nucleon decay
- Late 2013: Begin scintillator fill
  - Study backgrounds
- 2014: Add DBD isotope
- ~2017: Remove isotope, solar neutrino phase

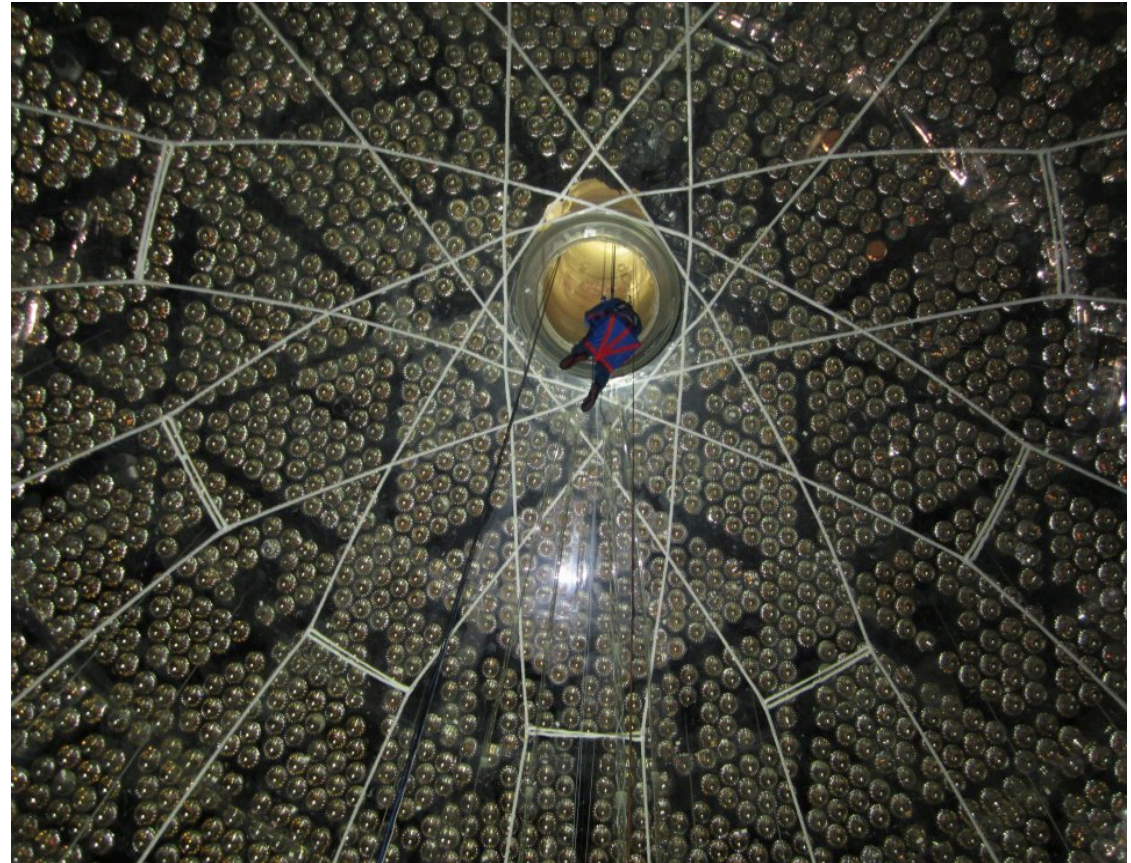
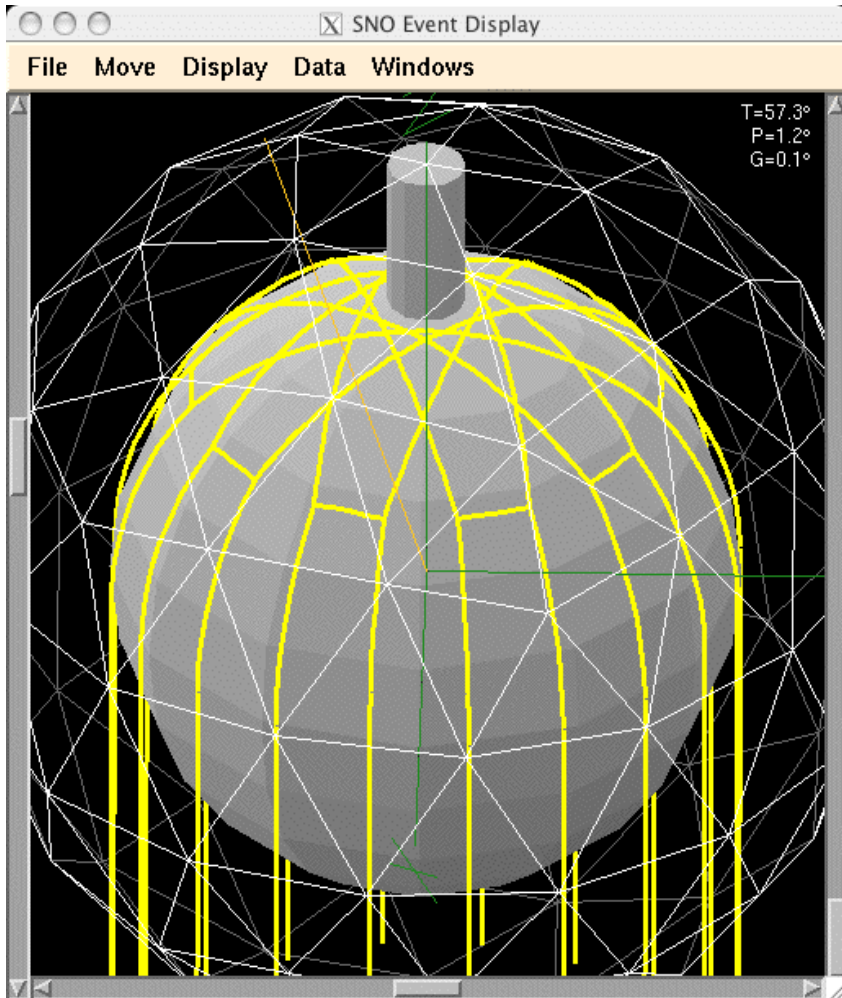
# Install AV Hold-Down Net



*Hold-down anchors and new floor liner installed*



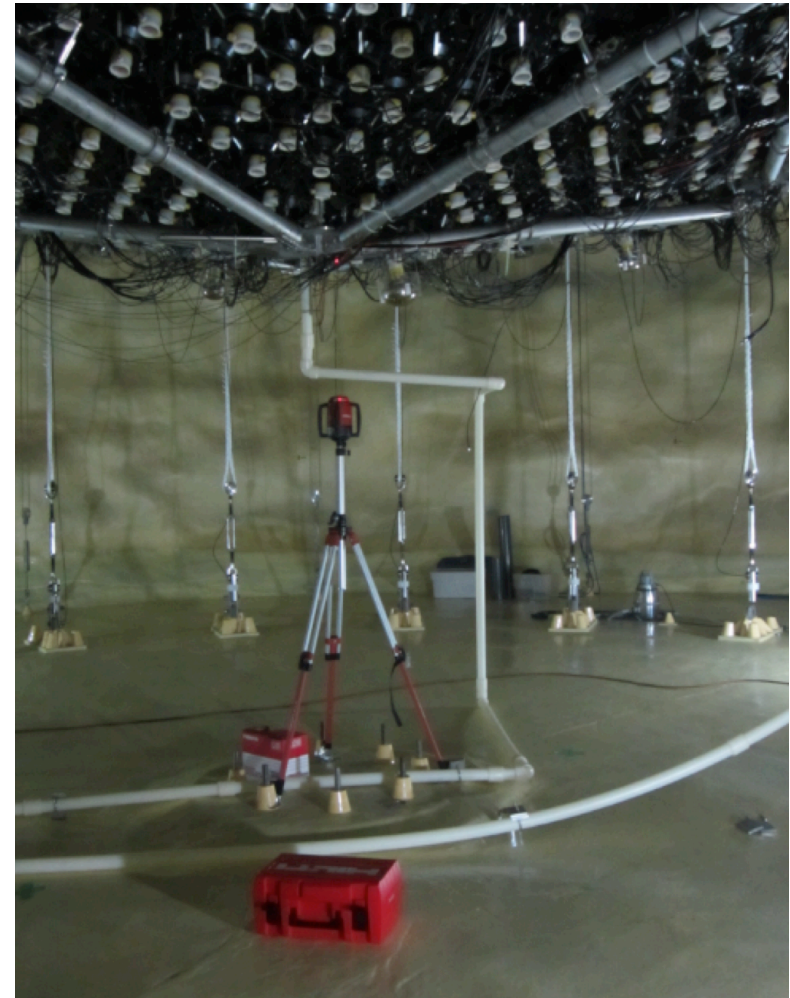
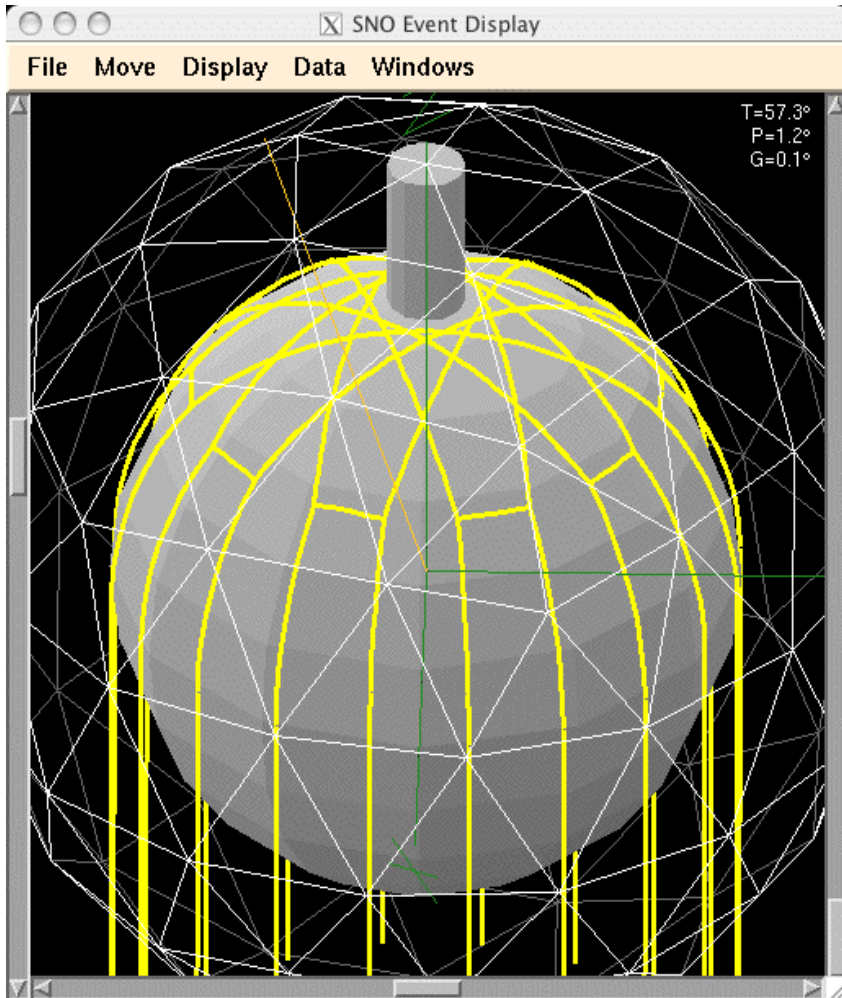
# Install AV Hold-Down Net



*Hold-down rope net installed and pre-tensioned*

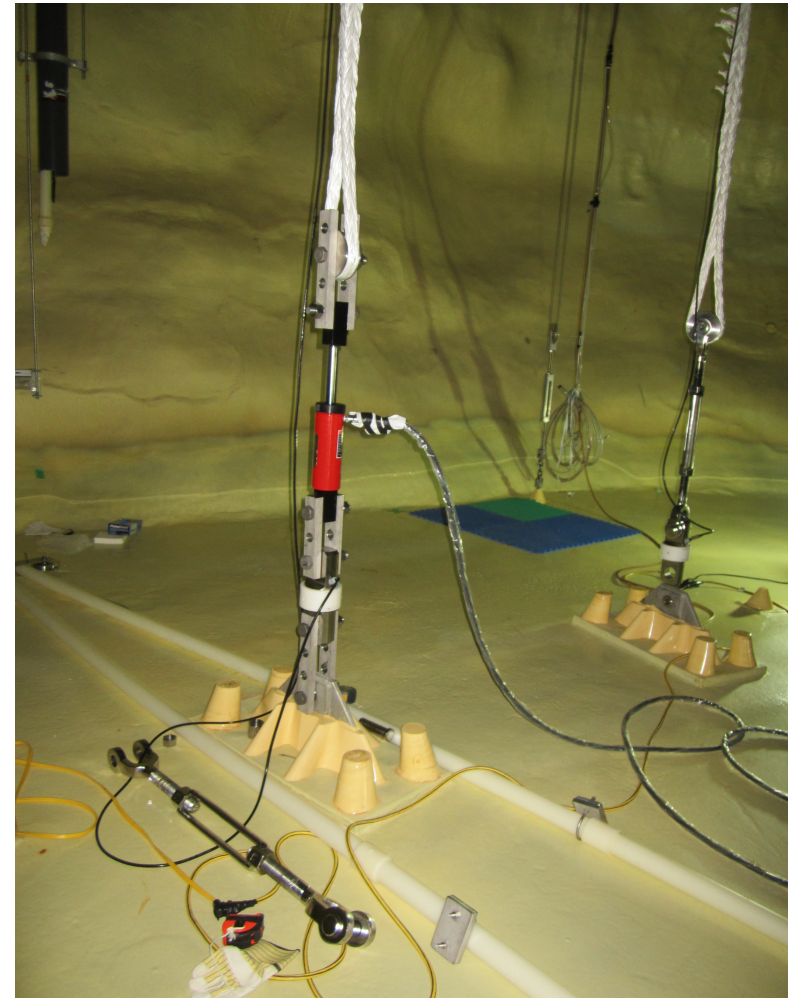
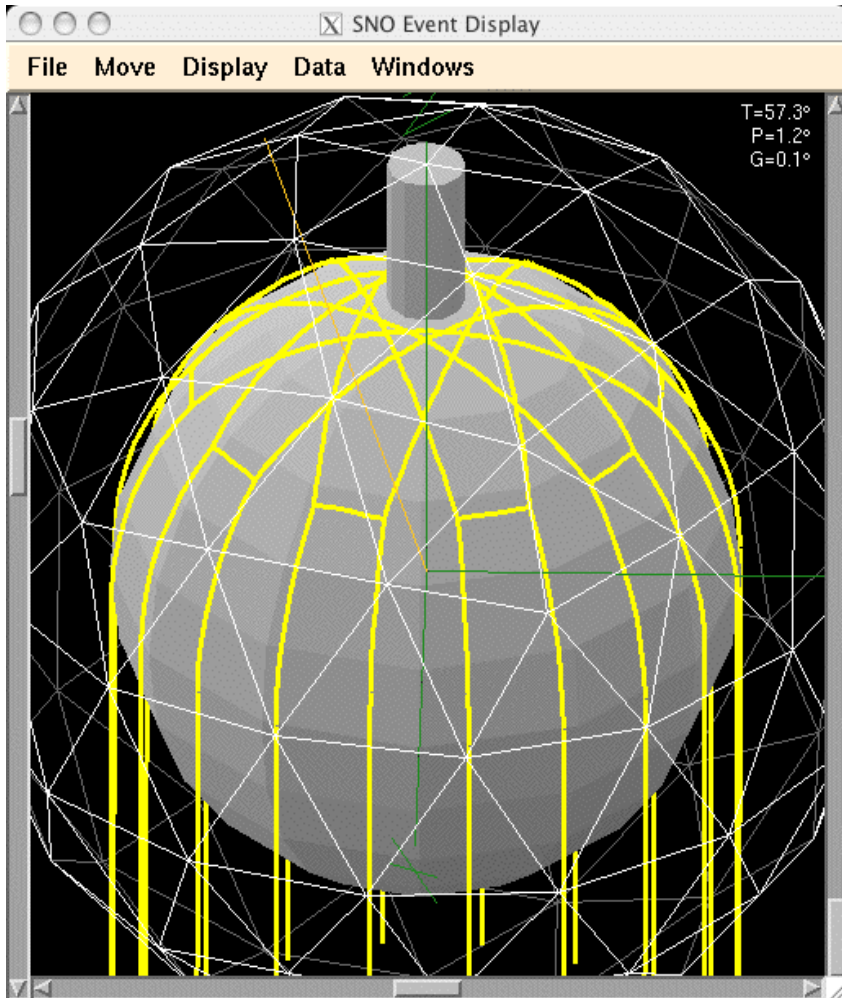


# Install AV Hold-Down Net



*Hold-down rope net installed and pre-tensioned*

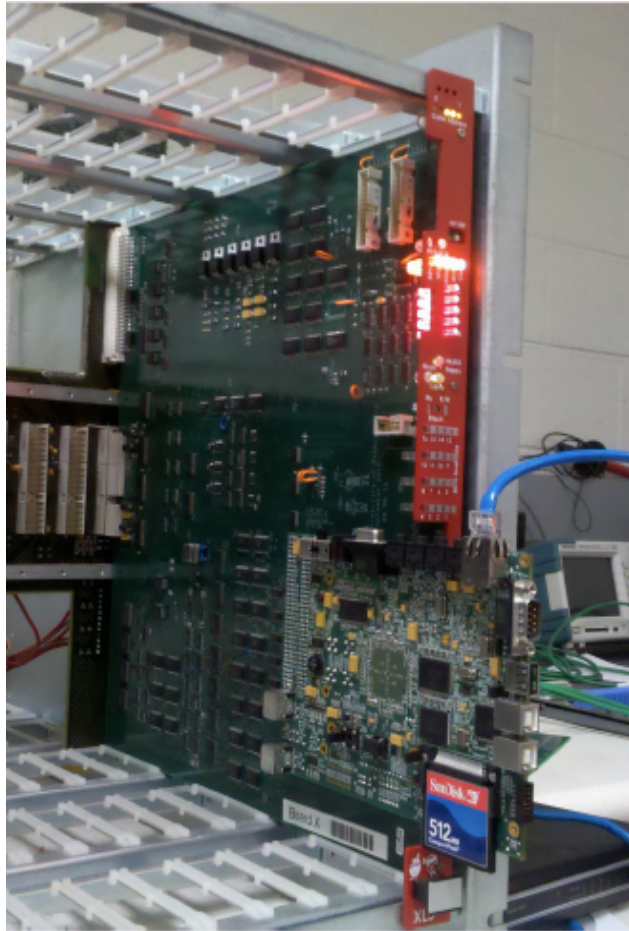
# Install AV Hold-Down Net



*Hold-down rope net installed and pre-tensioned*



# Electronics/DAQ Upgrades



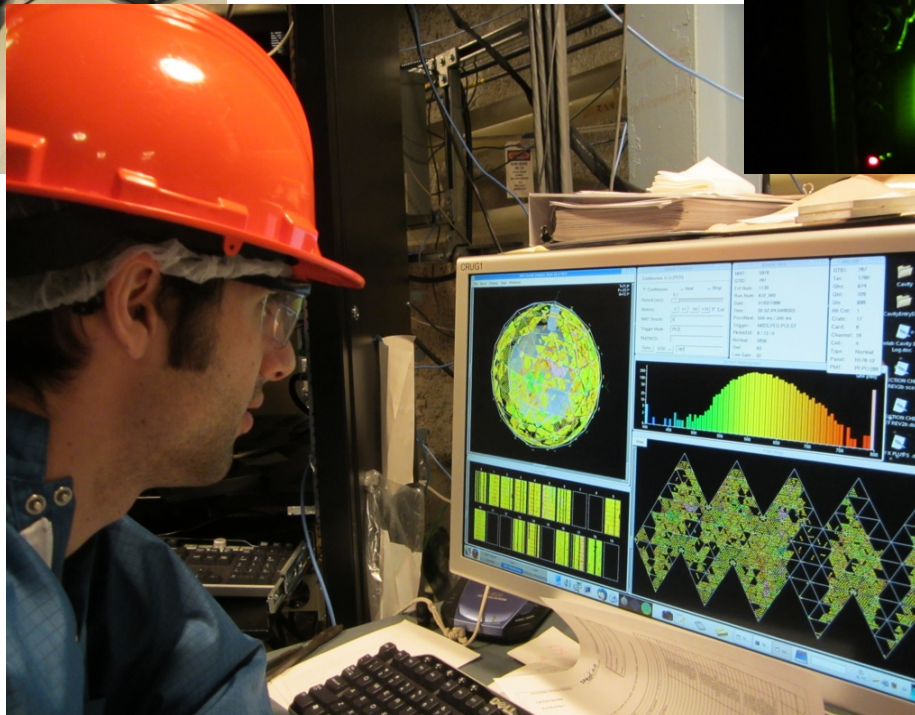
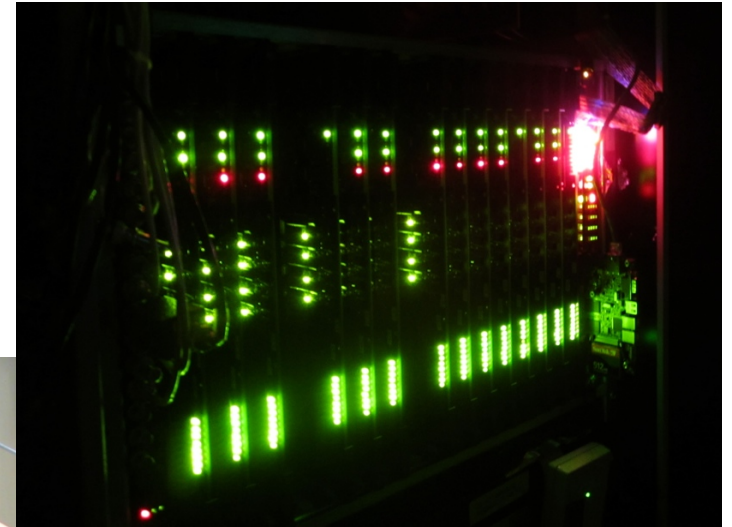
x19 in detector



x7 in detector

***Data rate ~100x greater than in SNO. Spring 2012: upgrade triggers, crate readout. Top rate >40 MByte/s***

# Electronics/DAQ Upgrades



***“Air Fill” running to test and integrate new DAQ and Electronics.***



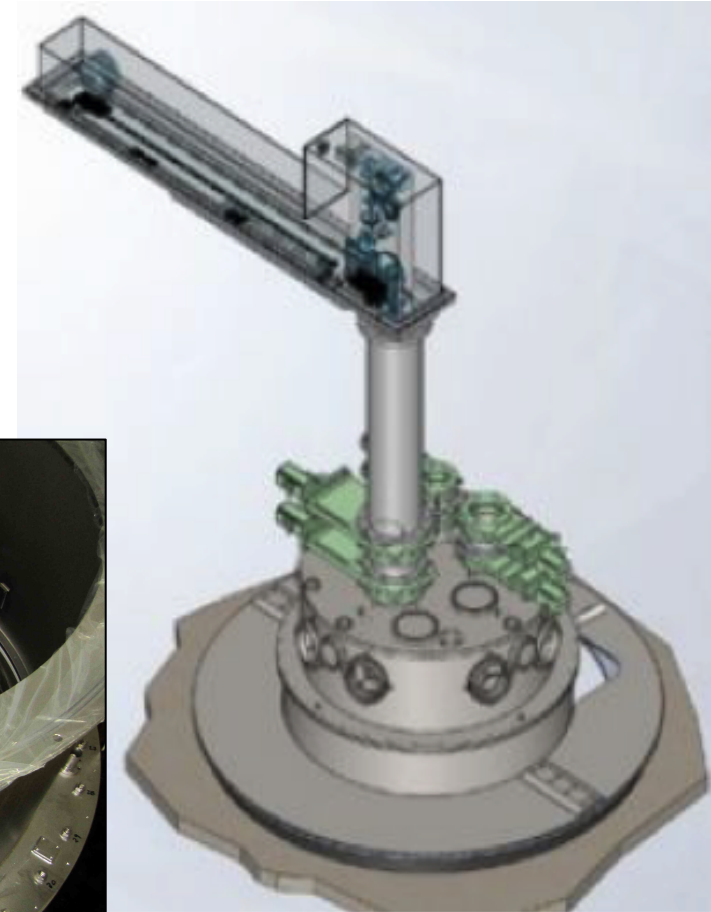
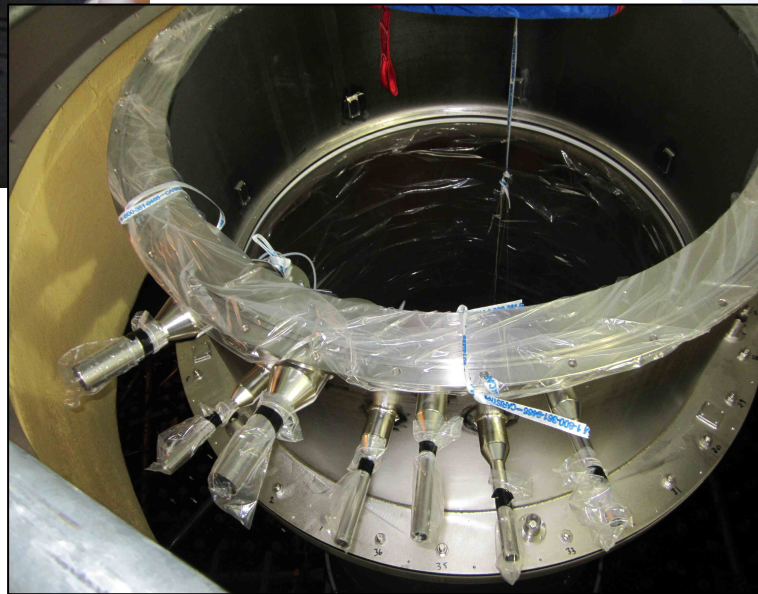
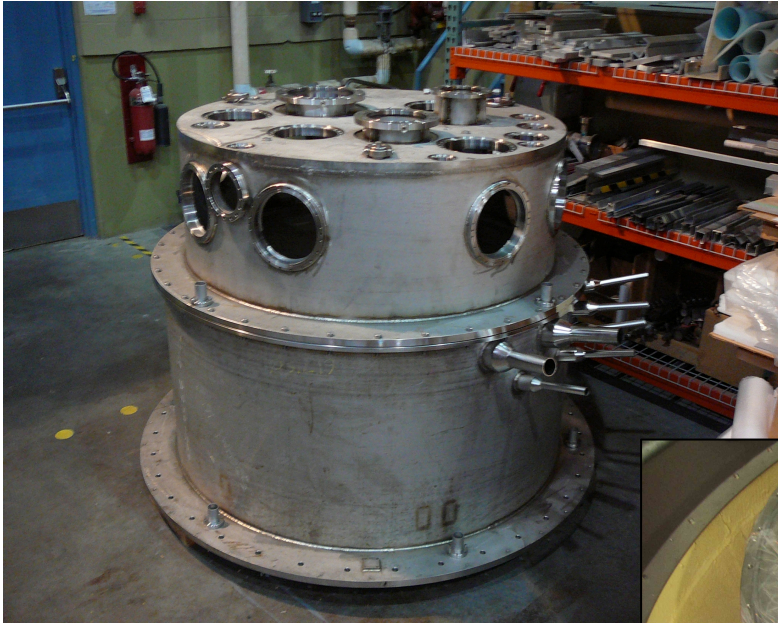
# Scintillator Process System



*All process vessels underground at SNOLAB, leak checking underway. Good progress on civil construction.*



# Calibration/Covergas System Upgrade



***Sealed covergas system with expandable bags for pressure equalization. Airlocks for calibration source deployment.***

# Calibration/Covergas System Upgrade

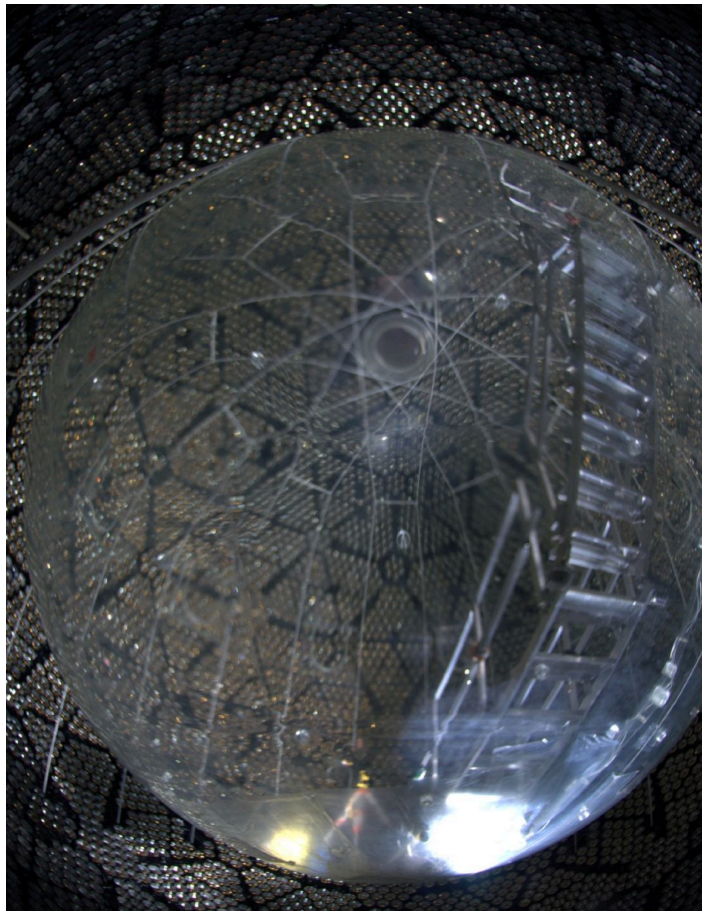
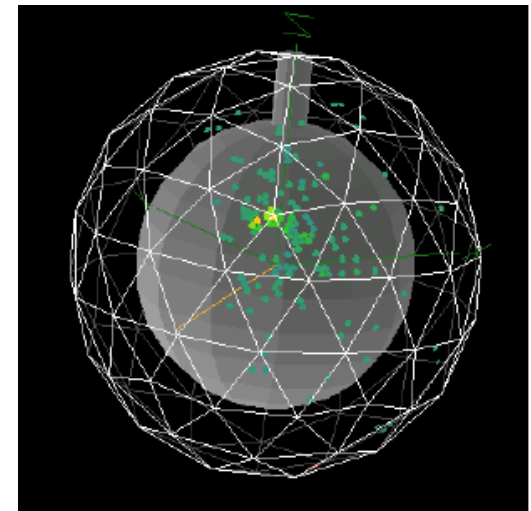
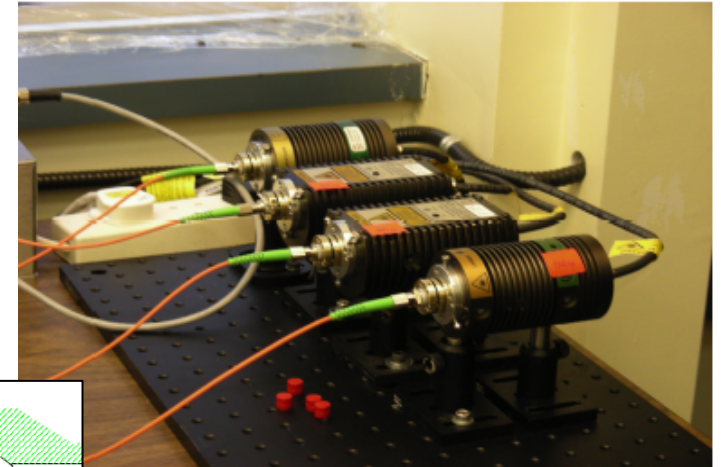
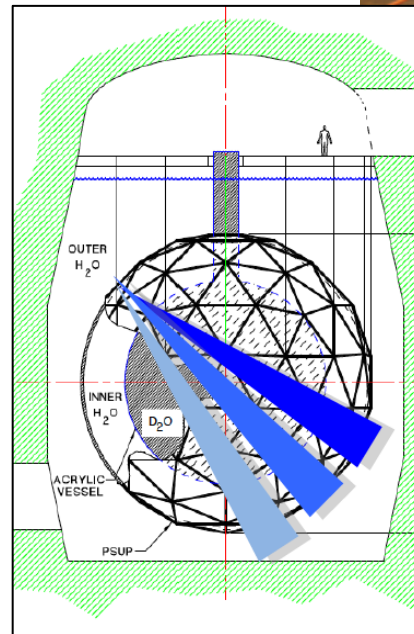


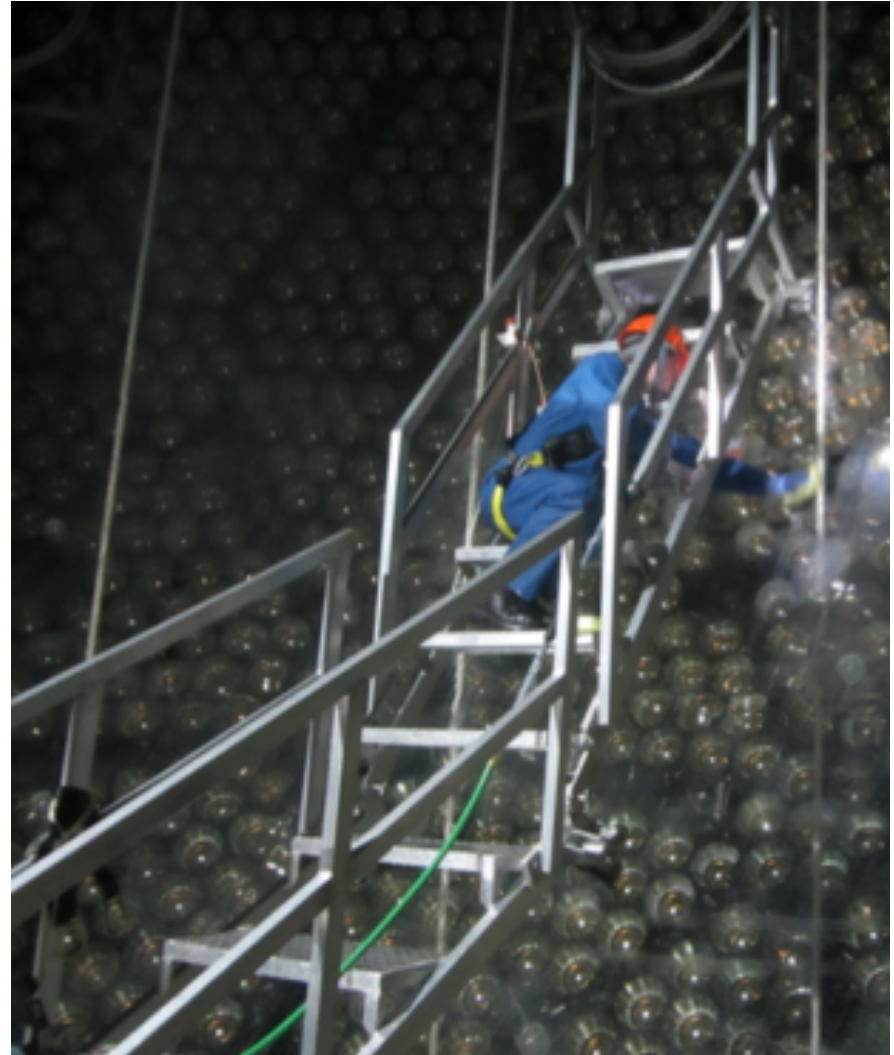
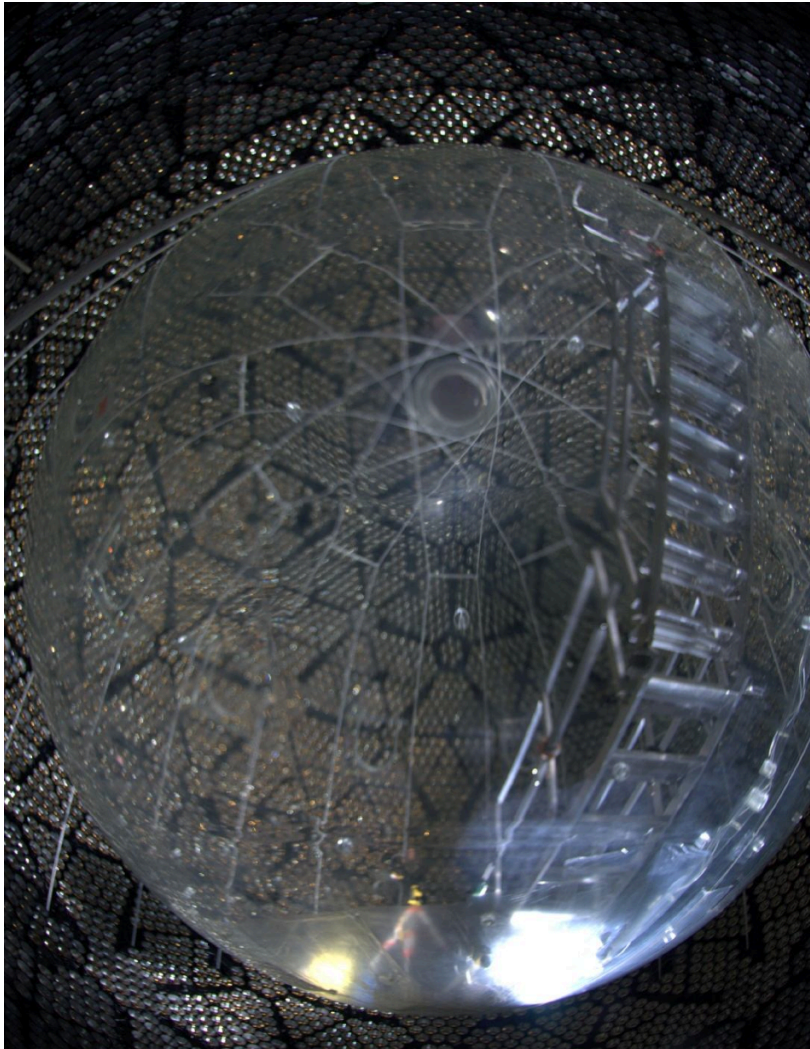
photo from newly installed  
calibration camera!



***Underwater cameras & lights for source position measurements  
& AV monitoring. Fiber optics for PMT calibrations.***



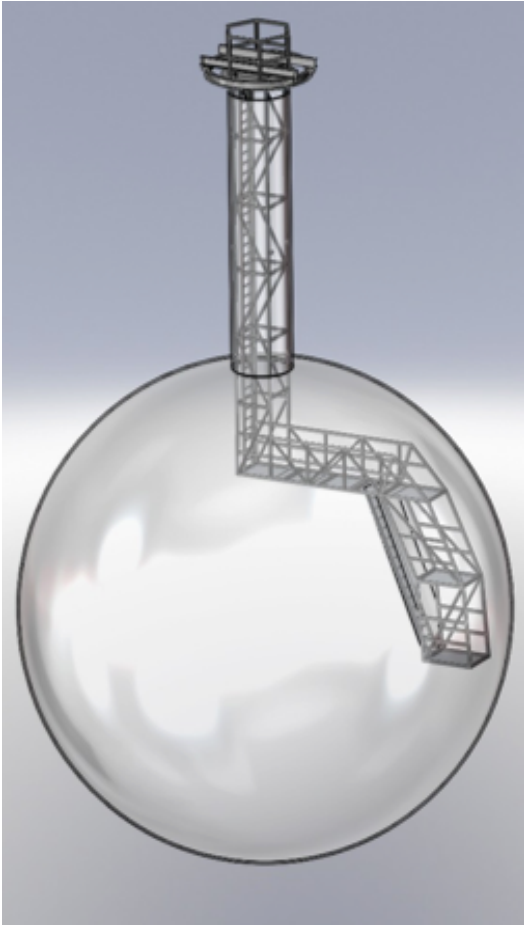
# Acrylic Vessel Cleaning



*Rotating ladder for lower hemisphere*



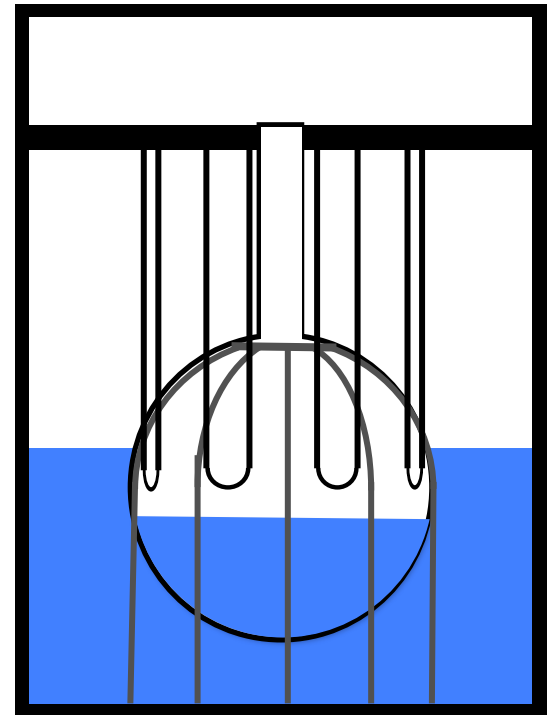
# Acrylic Vessel Cleaning



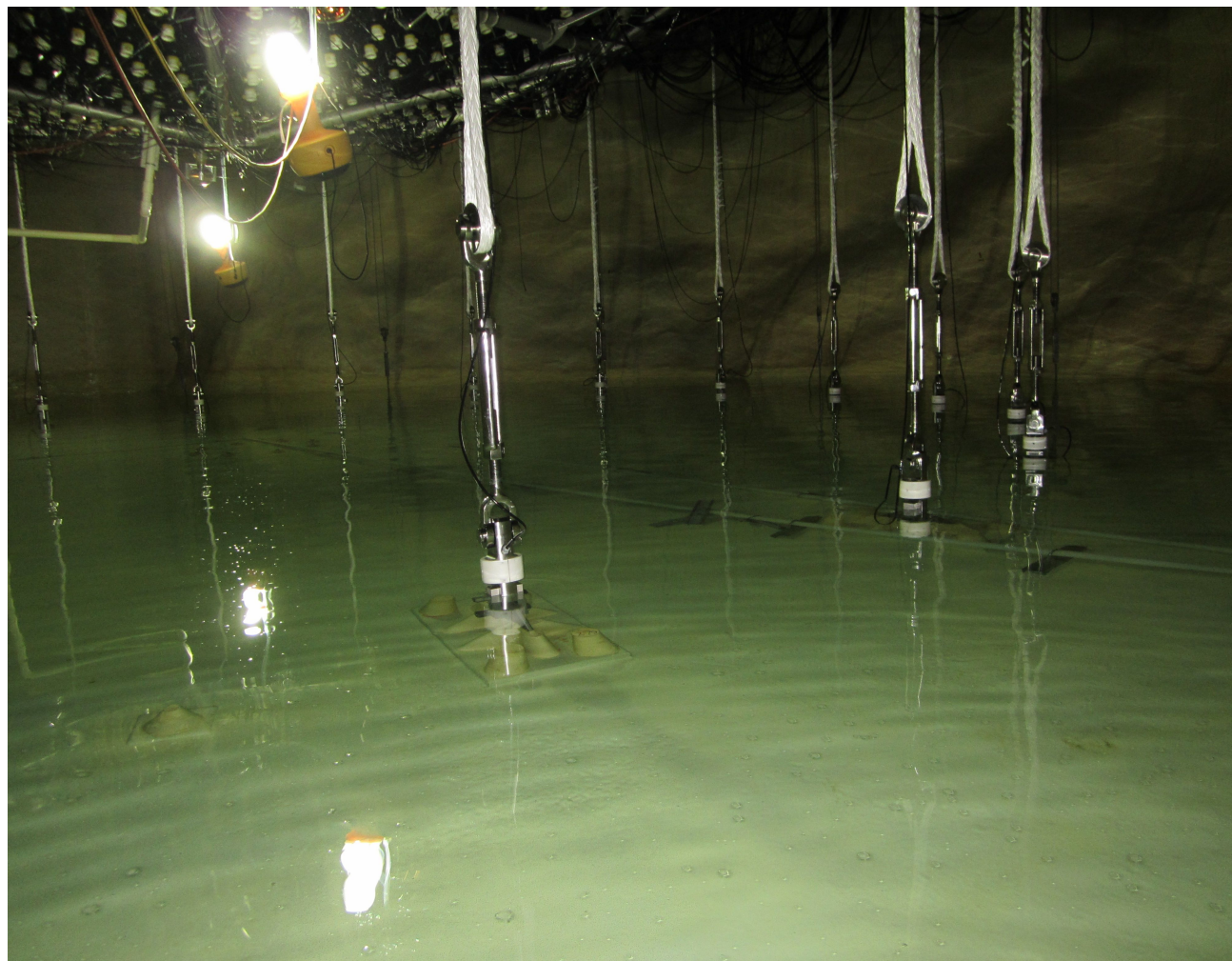
*Suspended platform to reach upper hemisphere*

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- Summer 2013: Begin water fill
  - Buoyant test of hold-down net
  - Study backgrounds and nucleon decay
- Late 2013: Begin scintillator fill
  - Study backgrounds
- 2014: Add DBD isotope
- ~2017: Remove isotope, solar neutrino phase



# Water Fill Underway



*Currently ~6' of water in the bottom of cavity. Hold this level to circulate and purify, check new floor liner for leaks.*

*Then, on to water fill, scintillator fill, and double beta decay!*