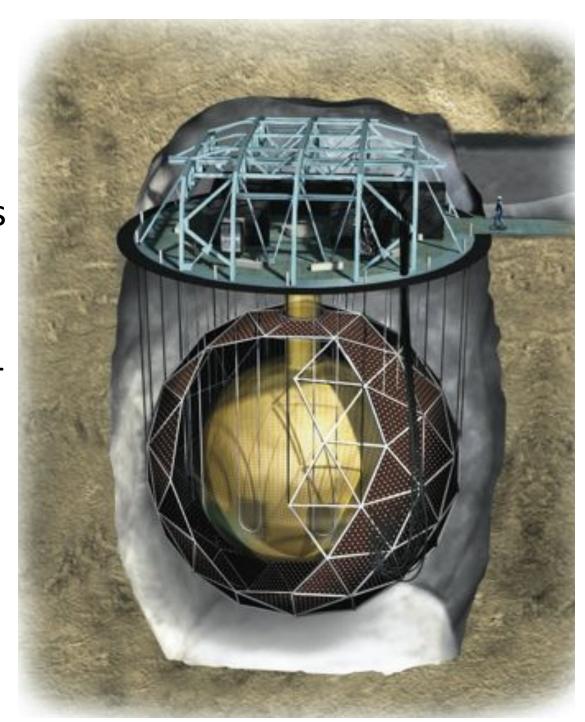


CAP Congress 2013, Montreal May 30th, 2013

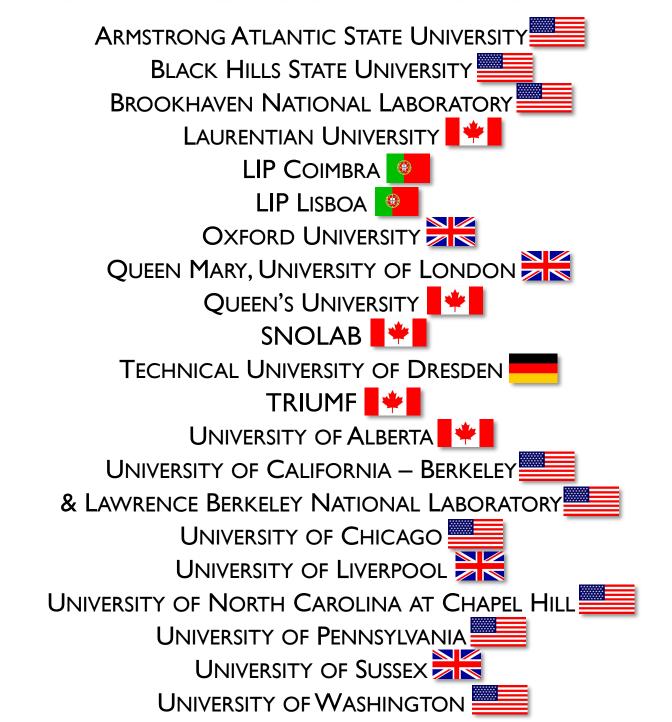
Alex Wright
IPP/Queen's University
For the SNO+ Collaboration



- SNO heavy water replaced by 780 tonnes of liquid scintillator
- ~9500 PMTs
- 1500 + 5300 tons ultrapure water shielding
- New rope net to hold down the 6m radius acrylic vessel
- 6800' underground in SNOLAB



The SNQ Collaboration



SNO+ Physics



Low Energy Solar Neutrinos

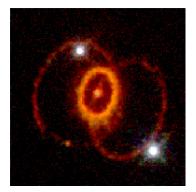




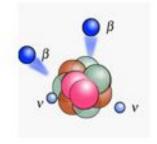


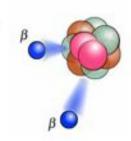
Geo-Neutrinos

Supernova Neutrinos



Neutrinoless Double Beta Decay





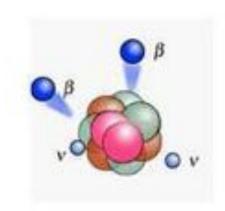
Neutrinoless Double Beta Decay

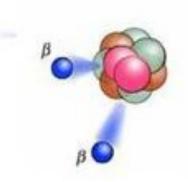
- Are neutrinos Majorana or Dirac particles?
 - Are they their own anti-particles?
- In double beta decay, a nucleus releases two electrons and two antineutrinos:

$$(A, Z) \rightarrow (A, Z + 2) + 2e^{-} + 2\overline{\nu}_{e}$$

 If neutrinos are Majorana, sometimes neutrinoless double beta decay occurs:

$$(A, Z) \rightarrow (A, Z + 2) + 2e^{-}$$





Detection of neutrinoless double beta decay proves that neutrinos are Majorana and provides information about the neutrino mass.

Searching for neutrinoless double beta decay involves looking for a tiny monoenergetic peak at the end of a large double beta decay continuum.

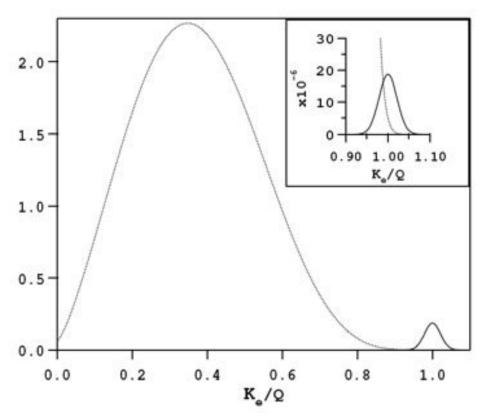
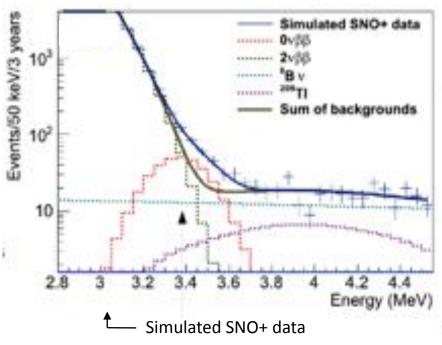


Image from Elliott and Vogel, hep-ph/0202254

D.B.D. experiments need good energy resolution, low backgrounds, and large amounts of isotope.

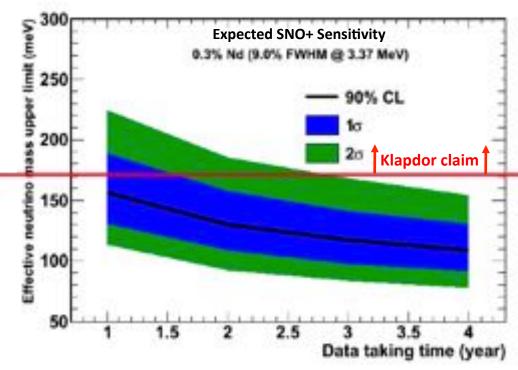
Neutrinoless Double Beta Decay in SNO+

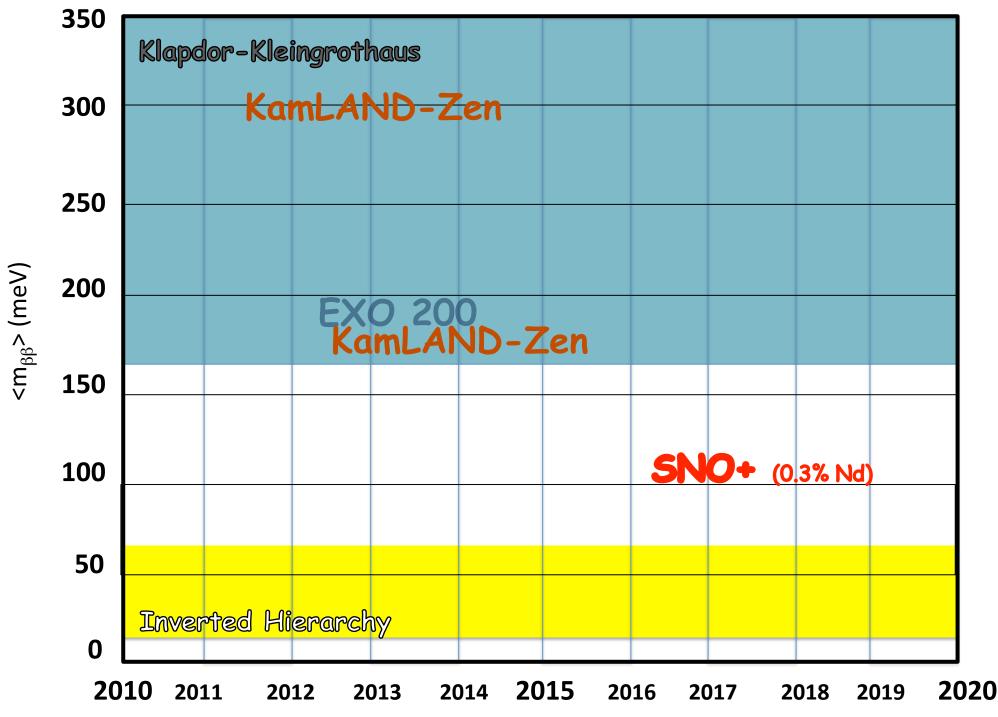
Loading neodymium into the SNO+ scintillator gives 140kg ¹⁵⁰Nd at 0.3% loading of natural neodymium (limited by optics).



- Signal at Klapdor level
- 2.4 live-years of data
- 50% fiducial volume
- Borexino-level backgrounds
 - ²¹⁴Bi 99.9% rejection
 - ²⁰⁸Tl 90% rejection





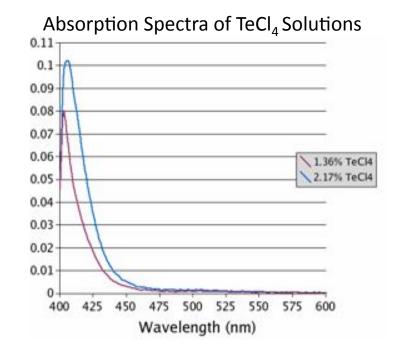


To improve the sensitivity of the SNO+ measurement, need more isotope.

$\beta\beta$ Isotopes with Q-values >2 MeV

isotope	Q-value [MeV]	natural abundance
⁴⁸ Ca	4.27	0.19%
¹⁵⁰ Nd	3.37	5.6%
⁹⁶ Zr	3.35	2.8%
¹⁰⁰ Mo	3.03	9.6%
⁸² Se	3.00	9.2%
¹¹⁶ Cd	2.80	7.5%
¹³⁰ Te	2.53	34.1%
¹³⁶ Xe	2.48	8.9%
¹²⁴ Sn	2.29	5.6%
⁷⁶ Ge	2.04	7.8%
¹¹⁰ Pd	2.01	11.8%

¹³⁰Te has highest natural abundance: 0.3% loading in SNO+ gives ~800kg of ¹³⁰Te isotope

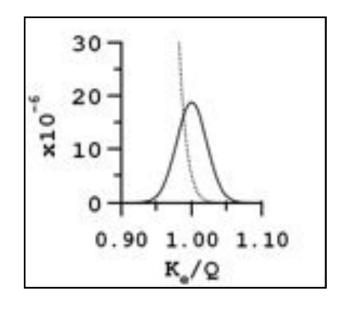


Te has no absorption lines, so no a priori limit on loading concentration...

$\beta\beta$ Isotopes with Q-values >2 MeV

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⁷⁶ Ge	2.04	7.8%
¹¹⁰ Pd	2.01	11.8%

In addition, the $0\nu\beta\beta/2\nu\beta\beta$ ratio is predicted to be about 100x higher in Te than Nd – this is an advantage given the limited energy resolution in scintillator.



Te-Loaded "Scintillators"

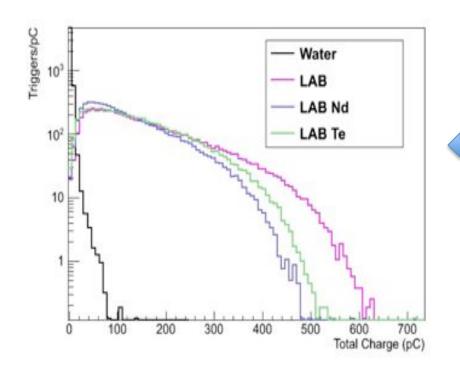


Te-Loaded Scintillator



SNO+ has developed stable Te-loaded LAB with good optical properties at percent-level loading.

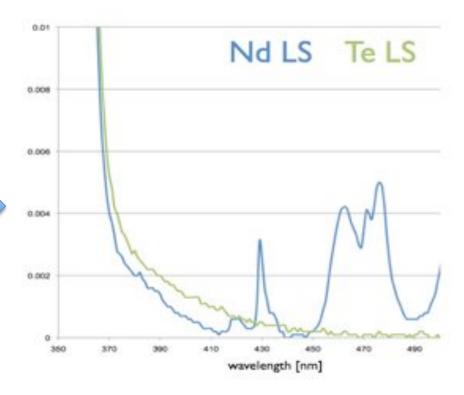
Te-Loaded Scintillator



Intrinsic light yield of Te-LS is good (better than Nd-LS)

 No absorption lines means that the light can be shifted out to 450-500nm

 Absorption may be further reduced by improvements in loading technique

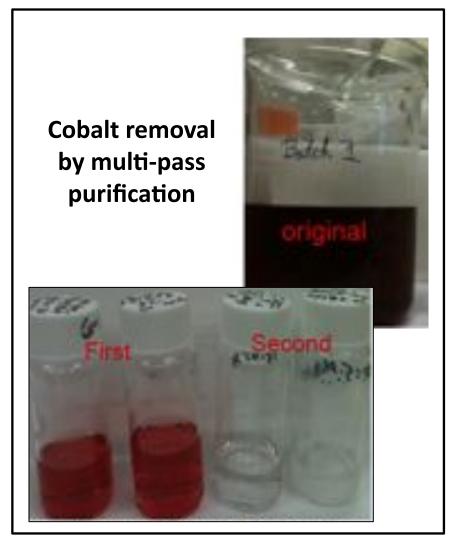


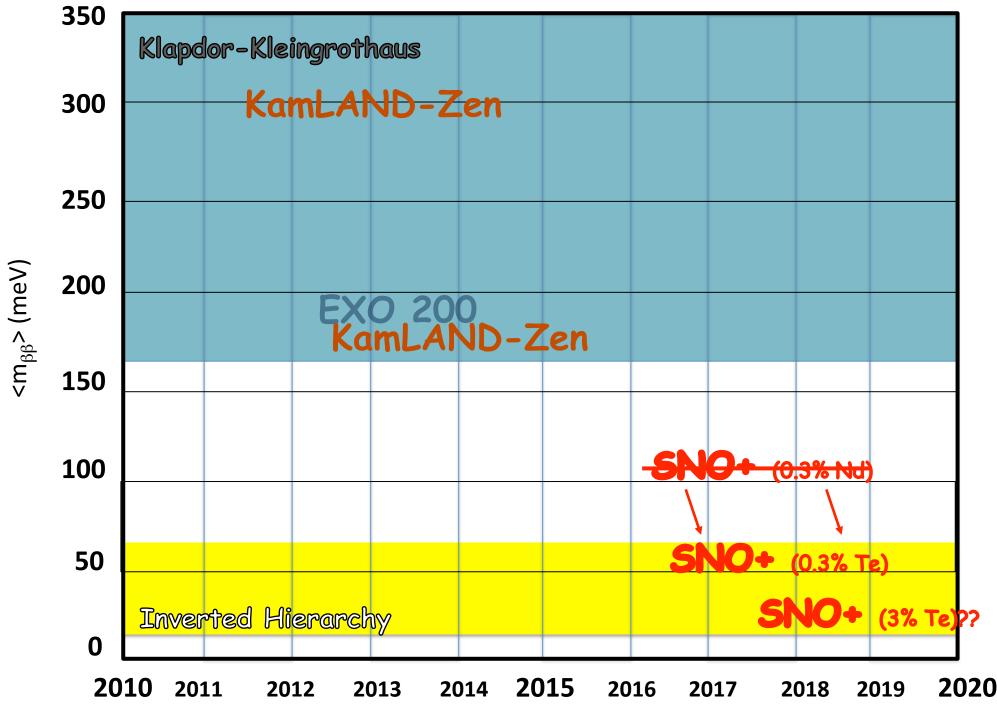
Backgrounds

- Lower energy (2.53 MeV) endpoint makes Te susceptible to more radiogenic backgrounds than Nd
- 2.6 MeV gamma from external ²⁰⁸Tl suppressed by fiducialization
 - Internal ²⁰⁸Tl at higher energy
- Te endpoint overlaps with ²¹⁴Bi spectrum (²³⁸U chain)
 - U-chain backgrounds in liquid scintillator can be extremely low (<2 decays/day/100T)
 - $-\ ^{214}Bi$ can be suppressed by more than a factor of 1000 using the 164µs ^{214}Bi ^{214}Po delayed coincidence
 - SNO+ has developed techniques to purify Te to acceptable U/ Th levels
 - Working to develop purification of other required chemicals

Backgrounds

- A number of isotopes that are cosmogenically produced from Te have longish half-lives and have decays that overlap the 0νββ energy region (214Sb, 126Sn, 88Y, 110Ag, etc...)
 - Effectively removed by our purification technique
 - Regrowth of cosmogenics requires at least a "polishing" purification underground





SNO+ will pursue Te deployment for double beta decay, starting with 0.3% target loading.

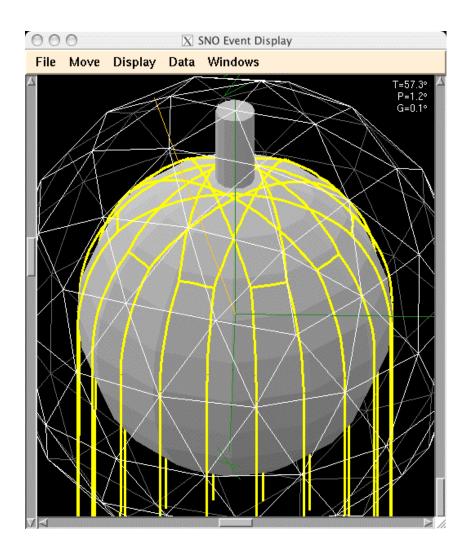
SNO+ Status and Schedule

- Current: Construction phase
 - Install AV hold-down net
 - Upgrade electronics/DAQ
 - Install scintillator purification system
 - Upgrade calibration/covergas system
 - Clean Acrylic Vessel
- Summer 2013: Begin water fill
 - Buoyant test of hold-down net
 - Study backgrounds and nucleon decay
- Late 2013: Begin scintillator fill
 - Study backgrounds
- 2014: Add DBD isotope
- ~2017: Remove isotope, solar neutrino phase



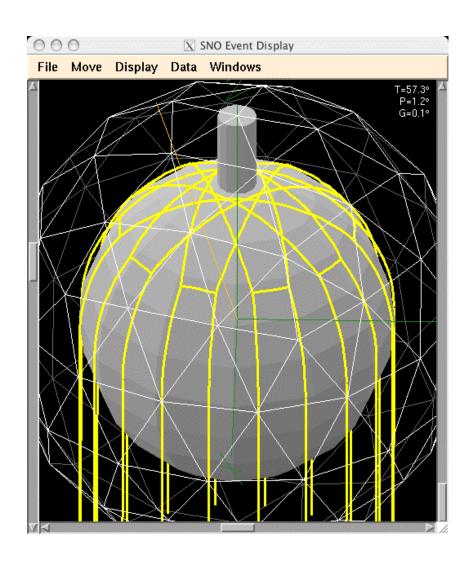


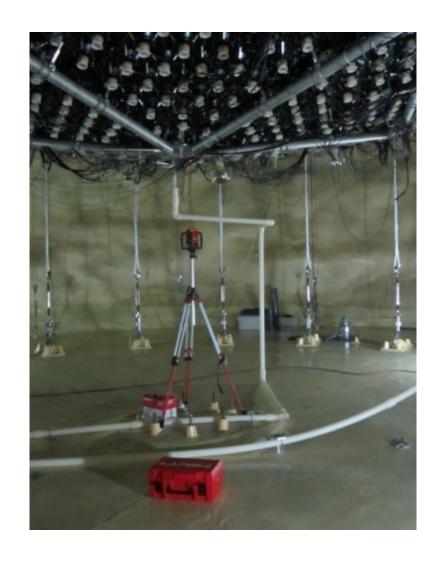
Hold-down anchors and new floor liner installed



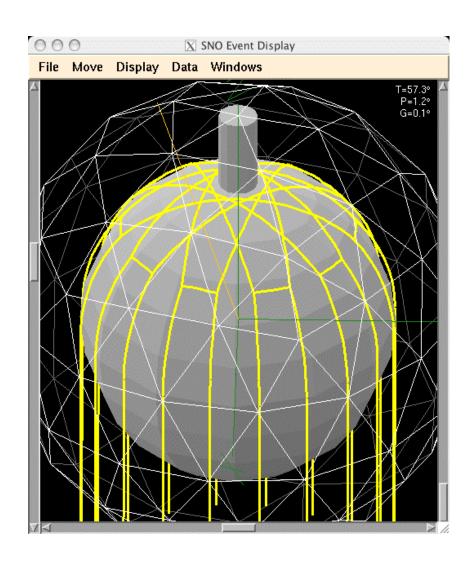


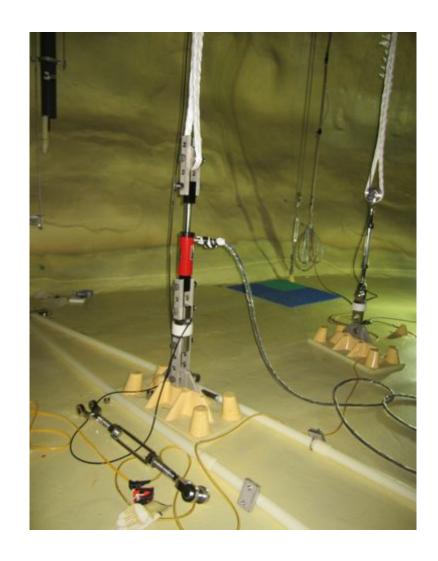
Hold-down rope net installed and pre-tensioned





Hold-down rope net installed and pre-tensioned





Hold-down rope net installed and pre-tensioned

Electronics/DAQ Upgrades



x7 in detector

x19 in detector

Data rate ~100x greater than in SNO. Spring 2012: upgrade triggers, crate readout. Top rate >40 MByte/s

Electronics/DAQ Upgrades



"Air Fill" running to test and integrate new DAQ and Electronics.

Scintillator Process System







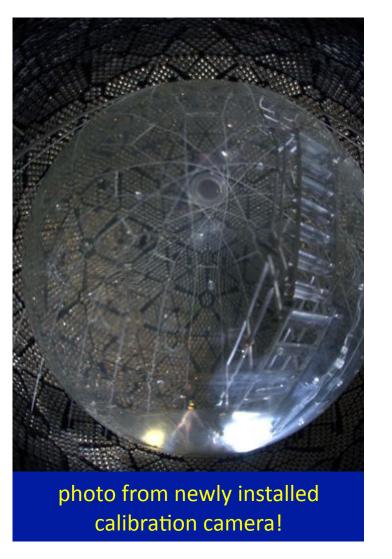
All process vessels underground at SNOLAB, leak checking underway. Good progress on civil construction.

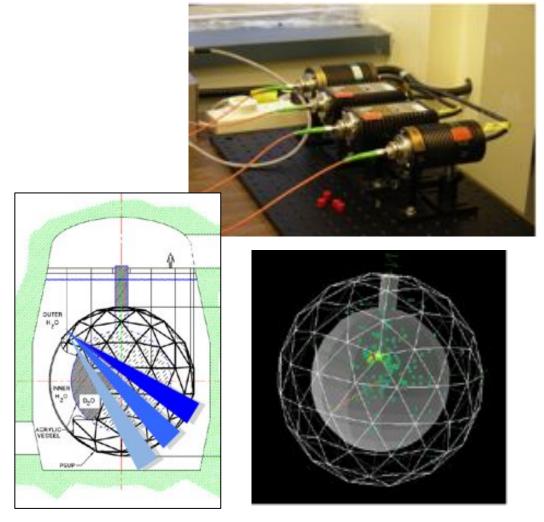
Calibration/Covergas System Upgrade



Sealed covergas system with expandable bags for pressure equalization. Airlocks for calibration source deployment.

Calibration/Covergas System Upgrade





Underwater cameras & lights for source position measurements & AV monitoring. Fiber optics for PMT calibrations.

Acrylic Vessel Cleaning





Rotating ladder for lower hemisphere

Acrylic Vessel Cleaning

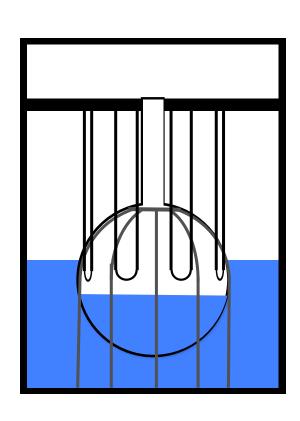




Suspended platform to reach upper hemisphere

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 - Study backgrounds and nucleon decay
- Late 2013: Begin scintillator fill
 - Study backgrounds
- 2014: Add DBD isotope
- ~2017: Remove isotope, solar neutrino phase



Water Fill Underway



Currently ~6' of water in the bottom of cavity. Hold this level to circulate and purify, check new floor liner for leaks.

Then, on to water fill, scintillator fill, and double beta decay!